# NEOARCHEAN BASEMENT, MANTLE ENRICHMENT AND CRUSTAL EXTRACTION IN CENTRAL ASIA: PETROGENESIS OF 2.5 GA AMPHIBOLITE AND METADIORITE IN NE CHINA 

HUICHUAN LIU*,**,†, JUN SHAO***, GUANGYOU ZHUs, and YINGLEI LIss,sss


#### Abstract

Archean basement in the Central Asian Orogenic Belt (CAOB) is relatively rare, but it has the potential to provide additional information on the processes of lithospheric mantle enrichment and crust extraction processes during the early history of the Earth. We identified Neoarchean amphibolite ( $2537-2565 \mathrm{Ma}$ ) and metadiorite $(2481-2539 \mathrm{Ma})$ in the Biliya area of the Erguna Terrane in the southeast CAOB. The amphibolite is geochemically MORB-like and has a weakly left-leaning REE pattern, and low zircon $\varepsilon \mathrm{Hf}(\mathrm{t})(-0.7-+6.2)$, and whole-rock $\varepsilon \mathrm{Nd}(\mathrm{t})(-1.7-+4.5)$ and $\varepsilon \mathrm{Hf}(\mathrm{t})(-1.9)$ values. Our petrogenetic modeling reveals that the amphibolite is derived from $\sim \mathbf{2 0} \%$ partial melting of the lithospheric mantle in the spinel stability field ( $\sim 65 \mathrm{~km}$ depth). The metadiorite shows near-zero $\varepsilon \mathrm{Nd}(\mathrm{t})(-0.5-+3.6)$ and $\varepsilon \mathrm{Hf}(\mathrm{t})(+0.5-+1.4)$ values and is likely derived from partial melting of mafic lower crust. The metadiorite and amphibolite likely formed in an extensional continental arc/back-arc setting and represent the Archean crystalline basement of the microcontinents within the CAOB. Three-staged mantle segregation and crust extraction processes have been proposed: (a) $20 \%$ melt extraction from primitive mantle-like lithospheric mantle, leaving behind a depleted mantle; (b) subduction-related fluid/melt metasomatism of the lithospheric mantle and its partial melting, generating the arc-type enriched mantle and mafic lower crust; and (c) partial remelting of the mafic lower crust produced the Tonalite-trondhjemite-granodiorite (TTG) crust.


Key words: Lithospheric mantle, Crust extraction, Continental basement, Neoarchean Earth, Central Asian Orogenic Belt

## INTRODUCTION

The Earth's lithospheric mantle shields the crust from the convecting asthenospheric mantle (Hofmann, 1988; van Hunen and Moyen, 2012). After four billion years of complex depletion and enrichment events, the present mantle is geochemically heterogeneous and dominated by four isotopic endmembers, namely the depleted mantle (DM), enriched mantle type I (EM I), enriched mantle type II (EM II) and high $\mu$ mantle (HIMU; Zindler and others, 1982; Anderson, 1983; White, 1985; Salters and Stracke, 2004). Although the nature of these four mantle endmembers has been clarified through extensive isotope studies, features of the lithospheric mantle and its mantle depletion and enrichment processes are still unclear. Furthermore, how the volume of continental crust was extracted from the upper mantle throughout the Earth history (crustal growth) remains poorly understood (Hawkesworth and Kemp, 2006; Kemp and others, 2009; Hawkesworth and others, 2010; Cawood and others, 2013; Cawood and Hawkesworth, 2019). The Archean is a

[^0]key stage of the mantle depletion and crustal extraction, and Archean rocks are important records of early Earth lithospheric mantle and crust (Hawkesworth and Kemp, 2006; Hawkesworth and others, 2010; Cawood and others, 2013; Cawood and Hawkesworth 2019).

The Central Asian Orogenic Belt (CAOB), situated between the Siberia and Baltica cratons to the north and the North China and Tarim cratons to the south, encompasses an immense region from the Urals in the west to the Russian Far East and includes tens of microcontinents and subduction-accretionary complexes (fig. 1A; Wilde and Zhou, 2015; Xiao and others, 2015; Zhou and others, 2018). The CAOB is known as the world's largest area of Phanerozoic juvenile crust formation, and much attention has been paid to its Phanerozoic crustal growth (Tang and others, 2010; Kovach and others, 2013; Eizenhöfer and others, 2015). No robust evidence of Archean upper mantle and crystalline basement have been reported for these microcontinents within the CAOB. Recent high-precision detrital zircon dating has revealed that sedimentary rocks previously ascribed to the Precambrian have Paleozoic or even Mesozoic maximum depositional ages (Miao and others, 2007; Xie and others, 2008; Wu and others, 2012; Xu and others, 2012; Sun and others, 2014; Zhang and others, 2014; Cui and others, 2015; Zhao and others, 2016; Zhou and others, 2018). Thus, the presence of Archean crystalline basement, the nature of its lithospheric mantle, and the thermo-tectonic processes that formed these microcontinents, are still unclear.

We collected amphibolite and metadiorite samples from drill-hole ZK6301 of the Biliya $\mathrm{Pb}-\mathrm{Zn}$ polymetallic deposit in the central-eastern Erguna Terrane (NE China; figs. 1B and 1C) and present data on their whole-rock geochemistry and $\mathrm{Sm}-\mathrm{Nd}$ and $\mathrm{Lu}-\mathrm{Hf}$ isotopes as well as zircon $\mathrm{U}-\mathrm{Pb}$ age and $\mathrm{Lu}-\mathrm{Hf}$ data. We discuss their magma source and the nature of their lithospheric mantle and crust extraction processes for these Neoarchean rocks of the CAOB.

## GEOLOGICAL BACKGROUND AND SAMPLING

The eastern CAOB contains the Stannovoy, Erguna, Xing'an and Burean-Jiamusi microcontinental terranes (fig. 1A). Previous research on the Precambrian basement of these microcontinents focused on inferred Precambrian strata (for example, IMBGMR, 1991). However, recent high-precision detrital zircon dating revealed that the so called "Precambrian" rocks show Paleozoic and even Mesozoic maximum depositional ages (for example, Miao and others, 2007; Xie and others, 2008; Wu and others, 2012; Xu and others, 2012; Sun and others, 2014; Zhang and others, 2014; Cui and others, 2015; Zhao and others, 2016; Zhou and others, 2018). Detrital zircon data from sedimentary units within the terranes reveals similar age peaks at $495 \mathrm{Ma}, 780 \mathrm{Ma}, 1825 \mathrm{Ma}$ and 2600 Ma to those reconized throughout the CAOB (Zhou and others, 2018). The eastern microcontinents within the CAOB share a common history to those identified elsewhere in the CAOB, including evidence for end Mesoproterozoic ( $\sim 1000 \mathrm{Ma}$ ) and early Paleozoic ( $\sim 500 \mathrm{Ma}$ ) orogenic events (Zhou and others, 2018). The Paleo-Asian Ocean dominated the Cambrian to Triassic tectonic evolution of the CAOB, and may have closed in the Late Permian to Early Triassic (Xiao and others, 2015). The tectonics of the eastern CAOB was also influenced by the Paleo-Pacific and Mongol-Okhotsk tectonic domains (Zhou and others, 2014; Wilde and Zhou, 2015). The Erguna Terrane is located between the Mongol-Okhotsk suture and Xinlin-Xiguitu-Toudaoqiao fault (fig. 1B) and is equivalent to the central Mongolia Terrane. Multiple Jurassic to Cretaceous volcanic and clastic units cover this terrane, including the Tamulangou, Manketouebo, Manitu, Baiyingaolao and Meiletu formations (fig. 2; IMBGMR, 1991). Permian to Cretaceous granitoids occur in a NE-SW trend across the terrane (fig. 1B).

The Derbuer ore field lies in central-eastern Erguna Terrane (NE China), and includes the De'rbuer, Erdaohezi and Jiawula-Chaganbuergen Ag-Pb-Zn polymetallic


Fig. 1. (A) Simplified tectonic map showing the microcontinents distributions in the CAOB (modified after Zhou and others, 2018). (B) Schematic geological map of NE China (modified after Wu and others, 2011). U-Pb age data are from Gou and others (2013), Luan and others, (2019) and Yang and others (2017). (C) Geological cross-section of the Biliya $\mathrm{Pb}-\mathrm{Zn}$ deposit in the Derbuer ore field.
deposits, Erentaolegai Ag deposit, and Biliya and Dongjun $\mathrm{Pb}-\mathrm{Zn}$ polymetallic deposits (Zhao and others, 2018). The Biliya $\mathrm{Pb-Zn}$ polymetallic deposit ( $50^{\circ} 59^{\prime} 51.9^{\prime \prime} \mathrm{N}, 120^{\circ}$ $57^{\prime} 36.5^{\prime \prime} \mathrm{E}$ ) is located in a 3.5 km long and 650 to 720 m wide, NW trending fault zone (Zhao and others, 2018). About 50 ore bodies are distrbuted along a strike of $295^{\circ}$ to $305^{\circ}$ (dip $65^{\circ}-85^{\circ}$; fig. 1C). Ore hosts are the volcanic rocks of the Tamulangou Formation (figs. 1C and 2), which comprise andesitic rocks (amygdaloidal/basaltic/ augite andesite), dacite and breccia tuff. The Tamulangou Fm. volcanic rocks yield a zircon U-Pb age of $141.6 \pm 1.8 \mathrm{Ma}$ (Ming and others, 2015), and the sphalerite in $\mathrm{Pb}-$ Zn ores were dated by $\mathrm{Rb}-\mathrm{Sr}$ at $141.6 \pm 1.9 \mathrm{Ma}$ (Zhao and others, 2017). Amphibolite


Fig. 2. Geological map showing the stratigraphic and igneous components of the Derbuer area (after IMBGMR, 1991).
and metadiorite were found in the four drill holes and they are in fault contact with the Tamulangou Formation (fig. 1C).

Six metadiorite (depth: 226 m ) and nine amphibolite (depth: 198 m ) samples were collected from the drill-hole ZK6301 at Biliya (figs. 1C and 2). The metadiorite samples comprise: (1) monzonitic two-mica schist (plagioclase + orthoclase + muscovite + biotite + quartz) with lepido-granoblastic texture (fig. 3A), (2) garnet two-mica quartz schist (plagioclase + muscovite + biotite + quartz + garnet) with porphyroblastic texture (fig. 3B), and (3) gneissic granodiorite with $\sim 30 \%$ andesine, $\sim 25 \%$ orthoclase, $\sim 15 \%$ quartz, $\sim 15 \%$ biotite, and $\sim 5 \%$ hornblende, along with the accessory minerals magnetite, zircon and apatite (figs. 3C and 3D). The amphibolite


Fig. 3. Microphotoes of the monzonitic mica schist (A), garnet two-mica quartz schist (B), gneissic granodiorite (C-D), and amphibolite (E-F). Abbrevations: Pl - plagioclase, Or - orthoclase, Bi - biotite, Ms - muscovite, Gr - garnet, Hbl - hornblende.
samples contain $\sim 75 \%$ hornblende, $\sim 20 \%$ plagioclase, $\sim 2 \%$ quartz and the accessory minerals chlorite, epidote, magnetite, zircon and apatite (figs. 3E and 3F).

## ANALYTICAL METHODS

## Zircon U-Pb Dating and in situ Zircon Lu-Hf Isotope

Zircon mineral separates were prepared by conventional heavy liquid and magnetic techniques. Grains were mounted in epoxy, polished and coated with gold and then photographed in transmitted and reflected light. Their internal texture was examined using cathodoluminescence (CL) imaging at the Institute of Geology, Chinese Academy of Geological Sciences (CAGS), Beijing.

SHRIMP zircon U-Pb dating: For sample D18-1-1, zircon U-Pb dating was conducted using a sensitive high-resolution ion microprobe (SHRIMP II) at the Beijing SHRIMP Center, Institute of Geology, CAGS. Cathodoluminescence (CL) images were made on zircons prior to U-Th- Pb analyses to reveal the internal textures and to guide the SHRIMP analyses. Detailed analytical procedures are described by Jian and others (2012). Common Pb correction used the 204 Pb method of Compston and others (1984). The SHRIMP analytical data presented are the mean values of five consecutive scans for each analytical spot. Errors quoted are at $1 \sigma$ level, whereas weighted mean ages of samples are quoted at $2 \sigma$ level.

LA-MC-ICPMS zircon U-Pb dating and in situ Lu-Hf isotopic analyses: For the four remaining samples (D18-2-1, D18-7-1, D18-3-1 and D18-6-1), U-Pb dating and trace element analysis of zircon were simultaneously conducted by LA-MC-ICPMS at the Wuhan Sample Solution Analytical Technology Co., Ltd., Wuhan, China. Detailed operating conditions for the laser ablation system and the ICP-MS instrument and data reduction are the same as description by Zong and others (2017). Zircon PLE and GJ-1 were used as external standards for U-Pb dating and trace element calibration, respectively. An Excel-based software ICPMSDataCal was used to perform off-line selection and integration of background and analyzed signals, time-drift correction and quantitative calibration for trace element analysis and U-Pb dating (Liu and others, 2010). Concordia diagrams and weighted mean calculations were made using Isoplot/Ex-ver2 (Ludwig, 2001). Zircon Lu-Hf isotopic analysis was carried out using
the LA-MC-ICPMS. All the settings yielded a signal intensity of $\sim 10$ Vat 180 Hf for the standard zircon 91500 with a recommended $176 \mathrm{Hf} / 177 \mathrm{Hf}$ ratio of $0.282293 \pm 28$ (Wu and others, 2006). Data were normalized to $176 \mathrm{Hf} / 177 \mathrm{Hf}=0.7325$, using exponential correction for mass bias. The ratio of $176 \mathrm{Yb} / 172 \mathrm{Yb}(0.5887)$ was also applied for the Yb correction.

## Whole-rock Element, Nd and Hf Isotope Determinations

Whole-rock major and trace elements: Whole rock samples for geochemistry were crushed to 200 -mesh using an agate mill for major and trace element analyses. The major oxides were analyzed by a wavelength X-ray fluorescence spectrometry at the Wuhan Sample Solution Analytical Technology Co., Ltd., Wuhan, China. Trace element analyses were also performed at the same company by a Perkin-Elmer Sciex ELAN 6000 ICP-MS. Detailed sample preparation and analytical procedure followed Li and others (2002).

Whole-rock Nd and Hf isotopes: All chemical preparations were performed on class 100 work benches within a class 1000 over-pressured clean laboratory. Column chemistry: Both the Nd and Hf isotope analyses were performed on a Neptune Plus MC-ICP-MS (Thermo Fisher Scientific, Dreieich, Germany) at the Wuhan Sample Solution Analytical Technology Co., Ltd, Hubei, China. Details of analytical methods are presented by Yang and others (2006). Normalizing factors used to correct the mass fractionation of Nd during the measurements were $146 \mathrm{Nd} / 144 \mathrm{Nd}=0.7219$ (Yang and others, 2005, 2006). For the Hf isotopes, mass discrimination correction was carried out via internal normalization to a $179 \mathrm{Hf} / 177 \mathrm{Hf}$ ratio of 0.7325 .

## ANALYTICAL RESULTS

## Geochronological and in situ Zircon Lu-Hf Isotopic Data

SHRIMP zircon U-Pb dating has been conducted on a metadiorite sample (D18-$1-1$ ) and LA-ICP-MS zircon U-Pb dating on two metadiorites samples (D18-2-1 and D18-7-1), and two amphibolite samples (D18-3-1 and D18-6-1). Detailed geochronological analytical results were listed in table 1. In situ zircon $\mathrm{Lu}-\mathrm{Hf}$ isotopic analyses were conducted on these five dated samples, with the results listed in table 2.

Sample D18-1-1: Thirty spot analyses were performed on 30 zircon grains from the metadiorite sample using SHRIMP. The 30 analyzed zircon crystals are mainly euhedral with length of 60 to $150 \mu \mathrm{~m}$ and aspect ratios ranging from 1:1 to 2:1 (fig. 4A). Most of the zircons contain well-defined, broad sector zones that become thinner from core to rim. They show high Th (47-560 ppm) and U (94-765 ppm) contents, and high $\mathrm{Th} / \mathrm{U}$ ratios ( $0.2-1.3$; fig. 4 G ). These CL images and high $\mathrm{Th} / \mathrm{U}$ ratios show typical features of igneous zircon (Hanchar and Miller, 1993). SHRIMP analyses on the magmatic cores plot on a discordia line with an upper concordia intercept age of $2481 \pm 21 \mathrm{Ma}(\operatorname{MSWD}=2.6, \mathrm{n}=30)$, which is consistent with the weighted mean $207 \mathrm{~Pb} / 206 \mathrm{~Pb}$ age of the most concordant analyses: $2494 \pm 61 \mathrm{Ma}$ (MSWD $=0.07, \mathrm{n}=$ 10 ; fig. 4A). This age ( $2481 \pm 21 \mathrm{Ma}$ ) is interpreted as the magma crystallization age of the protolith. Twenty zircons from sample D18-1-1 have a limited range of initial $176 \mathrm{Hf} / 177 \mathrm{Hf}$ ratios of 0.28118 to 0.28135 and $\varepsilon \mathrm{Hf}(\mathrm{t})$ values of -0.4 to +5.9 (fig. 5 A ). These values do not change with an increase in discordance, suggesting that the LuHf system remained closed when Pb loss occurred. The two-staged model ages are in the range of 2642 Ma to 3043 Ma (table 2).

Sample D18-2-1: Twenty analyses have been carried out on 20 zircon grains from this metadiorite sample. The twenty zircon crystals are euhedral to subhedral and some fragmentary, with length between 80 and $150 \mu \mathrm{~m}$, and contain broad cores and oscillatory thinner rims (fig. 4B). Our LA-ICPMS analyses concentrate in the cores and have high $\mathrm{Th} / \mathrm{U}$ ratios of 0.2 to 0.9 (fig. 4 G ). These characteristics indicate an
Table 1
Zircon $U-P b$ analytical results of the Biliya amphibolite (Samples D18-1-1, D18-2-1 and D18-7-1) and metadiorite

| Samples | $\mathrm{Th} / \mathrm{U}$ | ${ }^{207} \mathrm{~Pb} /{ }^{206} \mathrm{~Pb}$ | $2 \sigma$ | ${ }^{207} \mathrm{~Pb} /{ }^{235} \mathrm{U}$ | $2 \sigma$ | ${ }^{206} \mathrm{~Pb} /{ }^{238} \mathrm{U}$ | $2 \sigma$ | ${ }^{207} \mathrm{~Pb} /{ }^{206} \mathrm{~Pb}$ | $2 \sigma$ | ${ }^{207} \mathrm{~Pb} /{ }^{235} \mathrm{U}$ | $2 \sigma$ | ${ }^{206} \mathrm{~Pb} /{ }^{238} \mathrm{U}$ | $2 \sigma$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| D18-1-1-01 | 0.63 | 0.15780 | 0.00400 | 10.32630 | 0.30420 | 0.46760 | 0.01020 | 2432.4 | 43.7 | 2464.5 | 27.3 | 2473.0 | 44.8 |
| D18-1-1-02 | 0.71 | 0.16510 | 0.00360 | 6.67700 | 0.18800 | 0.28790 | 0.00580 | 2509.3 | 37.0 | 2069.6 | 24.9 | 1631.1 | 28.9 |
| D18-1-1-03 | 0.58 | 0.16340 | 0.00370 | 5.23320 | 0.16570 | 0.22870 | 0.00590 | 2491.1 | 37.2 | 1858.0 | 27.0 | 1327.6 | 31.2 |
| D18-1-1-04 | 0.60 | 0.18150 | 0.00460 | 5.05770 | 0.14940 | 0.19960 | 0.00480 | 2666.4 | 41.4 | 1829.0 | 25.1 | 1173.0 | 25.5 |
| D18-1-1-05 | 0.68 | 0.16690 | 0.00580 | 3.61160 | 0.20110 | 0.15220 | 0.00430 | 2526.9 | 53.2 | 1552.1 | 44.3 | 913.0 | 24.3 |
| D18-1-1-06 | 0.74 | 0.15130 | 0.00380 | 4.69460 | 0.12860 | 0.22140 | 0.00380 | 2360.8 | 42.9 | 1766.3 | 23.0 | 1289.5 | 20.2 |
| D18-1-1-07 | 0.83 | 0.16030 | 0.00360 | 10.55620 | 0.26580 | 0.47050 | 0.00790 | 2458.3 | 38.3 | 2484.9 | 23.4 | 2485.7 | 34.6 |
| D18-1-1-08 | 0.58 | 0.15940 | 0.00540 | 4.61510 | 0.15410 | 0.20990 | 0.00620 | 2449.1 | 56.8 | 1752.0 | 27.9 | 1228.2 | 32.9 |
| D18-1-1-09 | 0.43 | 0.16210 | 0.00450 | 5.23670 | 0.33070 | 0.23080 | 0.01340 | 2479.6 | 46.6 | 1858.6 | 53.8 | 1338.8 | 70.2 |
| D18-1-1-10 | 0.67 | 0.16280 | 0.00440 | 8.95660 | 0.26390 | 0.39330 | 0.00720 | 2484.9 | 46.8 | 2333.6 | 27.0 | 2138.3 | 33.5 |
| D18-1-1-11 | 0.43 | 0.15990 | 0.00420 | 6.83590 | 0.19090 | 0.30630 | 0.00540 | 2454.6 | 44.1 | 2090.4 | 24.8 | 1722.6 | 26.8 |
| D18-1-1-12 | 0.68 | 0.16050 | 0.00370 | 10.41260 | 0.24580 | 0.46480 | 0.00650 | 2460.8 | 38.9 | 2472.2 | 21.9 | 2460.9 | 28.8 |
| D18-1-1-13 | 0.70 | 0.15780 | 0.00370 | 5.61940 | 0.17350 | 0.25500 | 0.00610 | 2431.8 | 39.8 | 1919.1 | 26.7 | 1464.4 | 31.4 |
| D18-1-1-14 | 0.66 | 0.15790 | 0.00360 | 10.13370 | 0.24270 | 0.46100 | 0.00640 | 2433.0 | 38.4 | 2447.1 | 22.2 | 2444.1 | 28.4 |
| D18-1-1-15 | 0.58 | 0.15850 | 0.00430 | 6.69140 | 0.21790 | 0.30310 | 0.00650 | 2439.2 | 45.2 | 2071.5 | 28.8 | 1706.5 | 32.3 |
| D18-1-1-16 | 0.67 | 0.15860 | 0.00470 | 8.68080 | 0.26510 | 0.39440 | 0.00770 | 2442.6 | 55.1 | 2305.1 | 27.9 | 2143.1 | 35.7 |
| D18-1-1-17 | 0.50 | 0.16130 | 0.00430 | 5.01770 | 0.17470 | 0.22400 | 0.00610 | 2469.4 | 45.1 | 1822.3 | 29.5 | 1302.8 | 32.4 |
| D18-1-1-18 | 0.59 | 0.16300 | 0.00430 | 10.50710 | 0.30960 | 0.46510 | 0.01160 | 2486.7 | 44.4 | 2480.5 | 27.4 | 2461.9 | 51.1 |
| D18-1-1-19 | 0.84 | 0.21240 | 0.00560 | 7.05990 | 0.20670 | 0.23690 | 0.00310 | 2924.4 | 42.6 | 2119.0 | 26.1 | 1370.7 | 16.4 |
| D18-1-1-20 | 0.52 | 0.15760 | 0.00410 | 6.06840 | 0.22350 | 0.27540 | 0.00800 | 2431.5 | 43.1 | 1985.7 | 32.1 | 1568.2 | 40.2 |
| D18-1-1-21 | 0.68 | 0.16440 | 0.00420 | 10.74770 | 0.29020 | 0.47080 | 0.00890 | 2500.9 | 43.2 | 2501.6 | 25.1 | 2486.9 | 38.9 |
| D18-1-1-22 | 1.26 | 0.16300 | 0.00380 | 5.36160 | 0.16130 | 0.23630 | 0.00550 | 2486.7 | 38.9 | 1878.7 | 25.8 | 1367.5 | 28.7 |
| D18-1-1-23 | 0.80 | 0.17550 | 0.00400 | 12.16290 | 0.27840 | 0.49860 | 0.00630 | 2610.8 | 39.0 | 2617.1 | 21.5 | 2607.7 | 27.1 |
| D18-1-1-24 | 0.49 | 0.18290 | 0.00520 | 12.61810 | 0.39140 | 0.49660 | 0.00980 | 2679.9 | 46.6 | 2651.6 | 29.2 | 2599.2 | 42.1 |
| D18-1-1-25 | 1.09 | 0.18300 | 0.00530 | 12.67250 | 0.38530 | 0.49890 | 0.00830 | 2680.6 | 48.2 | 2655.6 | 28.7 | 2609.2 | 35.7 |
| D18-1-1-26 | 0.61 | 0.16110 | 0.00440 | 5.85240 | 0.18920 | 0.26010 | 0.00480 | 2477.8 | 51.1 | 1954.2 | 28.1 | 1490.6 | 24.4 |
| D18-1-1-27 | 0.73 | 0.16210 | 0.00390 | 6.48690 | 0.19290 | 0.28740 | 0.00570 | 2477.5 | 40.4 | 2044.1 | 26.2 | 1628.3 | 28.6 |
| D18-1-1-28 | 0.41 | 0.16440 | 0.00380 | 5.19380 | 0.24260 | 0.22760 | 0.00950 | 2501.9 | 38.9 | 1851.6 | 39.8 | 1321.7 | 50.1 |
| D18-1-1-29 | 0.19 | 0.17080 | 0.00410 | 3.77390 | 0.09490 | 0.15850 | 0.00200 | 2565.1 | 40.6 | 1587.2 | 20.2 | 948.3 | 11.2 |
| D18-1-1-30 | 0.62 | 0.16490 | 0.00420 | 9.91120 | 0.25310 | 0.43270 | 0.00610 | 2505.9 | 44.0 | 2426.6 | 23.6 | 2318.0 | 27.6 |
| D18-2-1-01 | 0.39 | 0.16021 | 0.00428 | 4.74800 | 0.14981 | 0.21228 | 0.00381 | 2457.7 | 44.6 | 1775.8 | 26.5 | 1241.0 | 20.3 |
| D18-2-1-02 | 0.59 | 0.16628 | 0.00394 | 11.12043 | 0.25426 | 0.48107 | 0.00368 | 2520.7 | 40.6 | 2533.3 | 21.4 | 2531.9 | 16.1 |
| D18-2-1-03 | 0.68 | 0.17532 | 0.00440 | 5.82485 | 0.14677 | 0.24011 | 0.00324 | 2609.0 | 40.9 | 1950.1 | 21.9 | 1387.3 | 16.9 |
| D18-2-1-04 | 0.63 | 0.16791 | 0.00354 | 10.99063 | 0.23694 | 0.47212 | 0.00488 | 2536.7 | 35.5 | 2522.3 | 20.1 | 2492.8 | 21.4 |

Table 1

| Samples | Th/U | ${ }^{207} \mathrm{~Pb}{ }^{206} \mathrm{~Pb}$ | $2 \sigma$ | ${ }^{207} \mathrm{~Pb} /{ }^{235} \mathrm{U}$ | $2 \sigma$ | ${ }^{206} \mathrm{~Pb}{ }^{238} \mathrm{U}$ | $2 \sigma$ | ${ }^{207} \mathrm{~Pb}{ }^{206} \mathrm{~Pb}$ | $2 \sigma$ | ${ }^{207} \mathrm{~Pb}{ }^{235} \mathrm{U}$ | $2 \sigma$ | ${ }^{206} \mathrm{~Pb} /{ }^{238} \mathrm{U}$ | $2 \sigma$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| D18-2-1-05 | 0.28 | 0.16242 | 0.00355 | 5.94299 | 0.13410 | 0.26416 | 0.00276 | 2481.2 | 37.0 | 1967.5 | 19.7 | 1511.1 | 14.1 |
| D18-2-1-06 | 0.45 | 0.26761 | 0.01338 | 13.23603 | 1.53039 | 0.28972 | 0.01833 | 3292.3 | 77.6 | 2696.6 | 109.2 | 1640.1 | 91.6 |
| D18-2-1-07 | 0.72 | 0.16461 | 0.00385 | 4.92059 | 0.12192 | 0.21604 | 0.00250 | 2503.4 | 38.7 | 1805.8 | 20.9 | 1260.9 | 13.2 |
| D18-2-1-08 | 0.76 | 0.16221 | 0.00333 | 8.62103 | 0.21351 | 0.38310 | 0.00540 | 2479.6 | 34.9 | 2298.8 | 22.6 | 2090.8 | 25.2 |
| D18-2-1-09 | 0.84 | 0.16124 | 0.00321 | 7.75578 | 0.16278 | 0.34745 | 0.00342 | 2468.8 | 33.6 | 2203.1 | 18.9 | 1922.4 | 16.4 |
| D18-2-1-10 | 0.92 | 0.16818 | 0.00311 | 11.65022 | 0.22107 | 0.50040 | 0.00443 | 2539.8 | 30.4 | 2576.7 | 17.8 | 2615.5 | 19.1 |
| D18-2-1-11 | 0.70 | 0.16866 | 0.00313 | 11.38570 | 0.22948 | 0.48748 | 0.00540 | 2544.1 | 31.2 | 2555.3 | 18.9 | 2559.8 | 23.4 |
| D18-2-1-12 | 0.67 | 0.16702 | 0.00332 | 12.09874 | 0.24501 | 0.52252 | 0.00437 | 2528.1 | 33.3 | 2612.1 | 19.1 | 2709.9 | 18.6 |
| D18-2-1-13 | 0.18 | 0.17534 | 0.00379 | 7.55803 | 0.18807 | 0.31133 | 0.00494 | 2609.6 | 41.5 | 2179.9 | 22.4 | 1747.2 | 24.3 |
| D18-2-1-14 | 0.86 | 0.16322 | 0.00347 | 11.32560 | 0.23083 | 0.50108 | 0.00417 | 2500.0 | 36.1 | 2550.3 | 19.1 | 2618.4 | 18.0 |
| D18-2-1-15 | 0.78 | 0.16770 | 0.00328 | 9.28395 | 0.28175 | 0.39878 | 0.00944 | 2534.9 | 33.5 | 2366.4 | 27.9 | 2163.4 | 43.5 |
| D18-2-1-16 | 0.63 | 0.17043 | 0.00318 | 12.76127 | 0.24443 | 0.54045 | 0.00490 | 2562.0 | 31.5 | 2662.2 | 18.1 | 2785.3 | 20.6 |
| D18-2-1-17 | 0.74 | 0.16190 | 0.00309 | 2.54135 | 0.06508 | 0.11336 | 0.00219 | 2475.6 | 33.2 | 1284.0 | 18.7 | 692.2 | 12.7 |
| D18-2-1-18 | 0.80 | 0.15689 | 0.00299 | 5.97266 | 0.12428 | 0.27503 | 0.00331 | 2433.3 | 31.9 | 1971.9 | 18.1 | 1566.3 | 16.7 |
| D18-2-1-19 | 0.49 | 0.15410 | 0.00320 | 4.68217 | 0.12304 | 0.21962 | 0.00407 | 2391.7 | 30.6 | 1764.1 | 22.0 | 1279.9 | 21.5 |
| D18-2-1-20 | 0.89 | 0.15405 | 0.00343 | 4.63909 | 0.10860 | 0.21779 | 0.00278 | 2391.7 | 38.0 | 1756.3 | 19.6 | 1270.2 | 14.7 |
| D18-7-1-01 | 0.71 | 0.17538 | 0.00671 | 12.11070 | 0.44124 | 0.49634 | 0.00741 | 2609.6 | 62.8 | 2613.0 | 34.2 | 2598.0 | 31.9 |
| D18-7-1-02 | 0.87 | 0.16582 | 0.00502 | 11.45901 | 0.33065 | 0.49356 | 0.00543 | 2515.7 | 50.9 | 2561.2 | 27.0 | 2586.1 | 23.5 |
| D18-7-1-03 | 0.71 | 0.15963 | 0.00480 | 6.42234 | 0.19576 | 0.28750 | 0.00444 | 2451.5 | 50.9 | 2035.3 | 26.8 | 1629.0 | 22.2 |
| D18-7-1-04 | 0.87 | 0.16885 | 0.00473 | 10.69111 | 0.31680 | 0.45163 | 0.00716 | 2546.0 | 46.9 | 2496.7 | 27.6 | 2402.5 | 31.8 |
| D18-7-1-05 | 0.90 | 0.17450 | 0.00533 | 9.14772 | 0.26225 | 0.37514 | 0.00419 | 2601.5 | 50.9 | 2352.9 | 26.3 | 2053.5 | 19.7 |
| D18-7-1-06 | 0.73 | 0.14888 | 0.00513 | 2.96908 | 0.14893 | 0.14182 | 0.00531 | 2333.0 | 59.3 | 1399.7 | 38.1 | 854.9 | 30.0 |
| D18-7-1-07 | 0.87 | 0.15891 | 0.00527 | 8.09416 | 0.28643 | 0.36427 | 0.00706 | 2444.1 | 56.2 | 2241.6 | 32.0 | 2002.4 | 33.4 |
| D18-7-1-08 | 0.59 | 0.17186 | 0.00542 | 11.91721 | 0.36800 | 0.49615 | 0.00558 | 2575.6 | 51.7 | 2597.9 | 29.0 | 2597.2 | 24.1 |
| D18-7-1-09 | 0.80 | 0.15522 | 0.00456 | 5.90184 | 0.17722 | 0.27303 | 0.00424 | 2405.6 | 50.0 | 1961.5 | 26.1 | 1556.2 | 21.5 |
| D18-7-1-10 | 0.75 | 0.16640 | 0.00458 | 12.06617 | 0.33647 | 0.52080 | 0.00646 | 2521.9 | 46.3 | 2609.6 | 26.2 | 2702.6 | 27.4 |
| D18-7-1-11 | 0.92 | 0.15982 | 0.00455 | 10.68882 | 0.28783 | 0.48101 | 0.00503 | 2454.0 | 48.1 | 2496.5 | 25.1 | 2531.6 | 22.0 |
| D18-7-1-12 | 0.64 | 0.16103 | 0.00475 | 10.49351 | 0.32050 | 0.46837 | 0.00651 | 2466.4 | 50.0 | 2479.3 | 28.4 | 2476.4 | 28.6 |
| D18-7-1-13 | 0.76 | 0.16667 | 0.00586 | 12.51679 | 0.43678 | 0.54171 | 0.00739 | 2524.4 | 59.4 | 2644.0 | 32.9 | 2790.6 | 30.9 |
| D18-7-1-14 | 0.69 | 0.17136 | 0.00562 | 12.30396 | 0.40946 | 0.51613 | 0.00684 | 2572.2 | 54.2 | 2627.9 | 31.3 | 2682.8 | 29.1 |
| D18-7-1-15 | 0.46 | 0.15816 | 0.00506 | 4.39960 | 0.15534 | 0.20005 | 0.00408 | 2436.1 | 59.7 | 1712.3 | 29.2 | 1175.6 | 21.9 |
| D18-7-1-16 | 0.72 | 0.16893 | 0.00498 | 9.20879 | 0.31282 | 0.39165 | 0.00847 | 2547.2 | 48.6 | 2359.0 | 31.2 | 2130.5 | 39.3 |
| D18-7-1-17 | 0.62 | 0.14966 | 0.00407 | 4.04436 | 0.14170 | 0.19216 | 0.00420 | 2342.3 | 46.6 | 1643.2 | 28.5 | 1133.1 | 22.7 |
| D18-7-1-18 | 0.70 | 0.16306 | 0.00469 | 9.94490 | 0.29124 | 0.43722 | 0.00708 | 2488.0 | 48.5 | 2429.7 | 27.1 | 2338.2 | 31.8 |
| D18-7-1-19 | 0.61 | 0.16095 | 0.00486 | 5.95097 | 0.18016 | 0.26465 | 0.00432 | 2465.7 | 51.4 | 1968.7 | 26.4 | 1513.6 | 22.0 |

Table 1
(continued)

| Samples | Th/U | ${ }^{207} \mathrm{~Pb} /{ }^{206} \mathrm{~Pb}$ | $2 \sigma$ | ${ }^{207} \mathrm{~Pb} /{ }^{235} \mathrm{U}$ | $2 \sigma$ | ${ }^{206} \mathrm{~Pb} /{ }^{238} \mathrm{U}$ | $2 \sigma$ | ${ }^{207} \mathrm{~Pb} /{ }^{206} \mathrm{~Pb}$ | $2 \sigma$ | ${ }^{207} \mathrm{~Pb} /{ }^{235} \mathrm{U}$ | $2 \sigma$ | ${ }^{206} \mathrm{~Pb}{ }^{238} \mathrm{U}$ | $2 \sigma$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| D18-7-1-20 | 0.69 | 0.15785 | 0.00502 | 8.23613 | 0.32448 | 0.37117 | 0.00979 | 2433.0 | 49.1 | 2257.3 | 35.7 | 2034.9 | 46.1 |
| D18-3-1-01 | 0.43 | 0.15901 | 0.00350 | 6.96429 | 0.16320 | 0.31668 | 0.00402 | 2455.6 | 37.0 | 2106.9 | 20.9 | 1773.5 | 19.7 |
| D18-3-1-02 | 0.54 | 0.16439 | 0.00328 | 11.23745 | 0.23722 | 0.49344 | 0.00518 | 2501.5 | 28.2 | 2543.0 | 19.8 | 2585.5 | 22.4 |
| D18-3-1-03 | 0.89 | 0.17176 | 0.00335 | 9.02528 | 0.18338 | 0.37895 | 0.00325 | 2575.9 | 32.7 | 2340.6 | 18.6 | 2071.4 | 15.2 |
| D18-3-1-04 | 0.85 | 0.16636 | 0.00348 | 9.13651 | 0.20712 | 0.39672 | 0.00509 | 2521.3 | 39.8 | 2351.8 | 20.8 | 2153.9 | 23.5 |
| D18-3-1-05 | 0.88 | 0.16744 | 0.00377 | 9.64299 | 0.23204 | 0.41631 | 0.00554 | 2532.4 | 37.7 | 2401.3 | 22.2 | 2243.7 | 25.3 |
| D18-3-1-06 | 0.59 | 0.16777 | 0.00392 | 10.43827 | 0.24822 | 0.44933 | 0.00502 | 2535.5 | 40.0 | 2474.5 | 22.1 | 2392.3 | 22.4 |
| D18-3-1-07 | 0.61 | 0.17058 | 0.00420 | 10.08269 | 0.25447 | 0.42720 | 0.00532 | 2564.8 | 42.1 | 2442.4 | 23.4 | 2293.1 | 24.1 |
| D18-3-1-08 | 0.90 | 0.16811 | 0.00356 | 7.75205 | 0.16494 | 0.33274 | 0.00302 | 2539.2 | 35.5 | 2202.7 | 19.2 | 1851.7 | 14.6 |
| D18-3-1-09 | 0.77 | 0.17316 | 0.00347 | 10.25113 | 0.21807 | 0.42766 | 0.00527 | 2588.6 | 33.3 | 2457.7 | 19.8 | 2295.1 | 23.8 |
| D18-3-1-10 | 0.96 | 0.17564 | 0.00358 | 9.39738 | 0.23258 | 0.38448 | 0.00508 | 2612.0 | 34.3 | 2377.6 | 22.8 | 2097.2 | 23.7 |
| D18-3-1-11 | 0.72 | 0.16822 | 0.00332 | 11.42763 | 0.23040 | 0.48966 | 0.00406 | 2539.8 | 33.0 | 2558.7 | 18.9 | 2569.2 | 17.6 |
| D18-3-1-12 | 0.82 | 0.16744 | 0.00379 | 7.86147 | 0.21861 | 0.33995 | 0.00696 | 2532.4 | 38.0 | 2215.3 | 25.1 | 1886.4 | 33.5 |
| D18-3-1-13 | 0.63 | 0.16775 | 0.00387 | 12.19161 | 0.29197 | 0.52320 | 0.00510 | 2535.5 | 38.9 | 2619.3 | 22.5 | 2712.7 | 21.6 |
| D18-3-1-14 | 0.60 | 0.16568 | 0.00402 | 7.17665 | 0.18260 | 0.31190 | 0.00360 | 2514.5 | 40.7 | 2133.6 | 22.7 | 1750.0 | 17.7 |
| D18-3-1-15 | 0.66 | 0.16438 | 0.00372 | 11.43358 | 0.28733 | 0.50029 | 0.00633 | 2500.9 | 38.0 | 2559.2 | 23.5 | 2615.0 | 27.2 |
| D18-3-1-16 | 0.86 | 0.16494 | 0.00368 | 11.29547 | 0.24869 | 0.49359 | 0.00426 | 2507.1 | 37.7 | 2547.8 | 20.6 | 2586.2 | 18.5 |
| D18-3-1-17 | 0.75 | 0.16745 | 0.00333 | 12.30614 | 0.24884 | 0.52920 | 0.00465 | 2532.4 | 33.3 | 2628.0 | 19.1 | 2738.1 | 19.7 |
| D18-3-1-18 | 0.78 | 0.16344 | 0.00330 | 11.56832 | 0.22701 | 0.51010 | 0.00399 | 2491.7 | 34.0 | 2570.1 | 18.4 | 2657.1 | 17.1 |
| D18-3-1-19 | 1.23 | 0.16438 | 0.00317 | 12.31539 | 0.23030 | 0.53988 | 0.00435 | 2500.9 | 32.4 | 2628.7 | 17.7 | 2782.9 | 18.3 |
| D18-3-1-20 | 0.90 | 0.16529 | 0.00335 | 9.73832 | 0.19959 | 0.42451 | 0.00417 | 2510.8 | 34.0 | 2410.3 | 18.9 | 2280.9 | 18.9 |
| D18-6-1-01 | 0.65 | 0.16152 | 0.00464 | 8.85085 | 0.28556 | 0.39202 | 0.00856 | 2471.3 | 43.5 | 2322.7 | 29.5 | 2132.2 | 39.6 |
| D18-6-1-02 | 0.73 | 0.16633 | 0.00435 | 11.95478 | 0.30806 | 0.51319 | 0.00566 | 2521.3 | 44.1 | 2600.9 | 24.2 | 2670.2 | 24.2 |
| D18-6-1-03 | 0.68 | 0.17200 | 0.00468 | 12.79324 | 0.33815 | 0.53251 | 0.00587 | 2577.5 | 45.4 | 2664.5 | 25.0 | 2752.0 | 24.7 |
| D18-6-1-04 | 0.74 | 0.16841 | 0.00460 | 8.13806 | 0.30182 | 0.34440 | 0.00892 | 2541.7 | 45.7 | 2246.5 | 33.6 | 1907.8 | 42.8 |
| D18-6-1-05 | 0.54 | 0.16597 | 0.00471 | 6.32709 | 0.18472 | 0.27305 | 0.00365 | 2517.6 | 47.8 | 2022.2 | 25.6 | 1556.3 | 18.5 |
| D18-6-1-06 | 0.53 | 0.17403 | 0.00541 | 9.62817 | 0.32379 | 0.39681 | 0.00723 | 2598.2 | 51.9 | 2399.9 | 31.0 | 2154.3 | 33.4 |
| D18-6-1-07 | 0.67 | 0.17923 | 0.00580 | 9.72224 | 0.30788 | 0.39060 | 0.00546 | 2646.0 | 53.7 | 2408.8 | 29.2 | 2125.6 | 25.4 |
| D18-6-1-08 | 0.72 | 0.17270 | 0.00539 | 10.41909 | 0.32171 | 0.43459 | 0.00630 | 2584.3 | 51.9 | 2472.7 | 28.7 | 2326.4 | 28.3 |
| D18-6-1-09 | 0.56 | 0.16704 | 0.00481 | 10.73094 | 0.31281 | 0.46189 | 0.00568 | 2528.1 | 48.2 | 2500.1 | 27.1 | 2447.9 | 25.1 |
| D18-6-1-10 | 0.79 | 0.17453 | 0.00491 | 11.35148 | 0.31167 | 0.46789 | 0.00463 | 2601.5 | 46.9 | 2552.4 | 25.7 | 2474.3 | 20.4 |
| D18-6-1-11 | 0.60 | 0.17531 | 0.00551 | 12.75297 | 0.38977 | 0.52525 | 0.00676 | 2609.0 | 46.8 | 2661.6 | 28.8 | 2721.4 | 28.6 |
| D18-6-1-12 | 0.62 | 0.16877 | 0.00563 | 8.52906 | 0.29507 | 0.36402 | 0.00614 | 2545.4 | 61.3 | 2289.0 | 31.5 | 2001.2 | 29.1 |
| D18-6-1-13 | 0.90 | 0.16785 | 0.00576 | 11.67735 | 0.40795 | 0.49985 | 0.00610 | 2536.1 | 58.5 | 2578.9 | 32.7 | 2613.2 | 26.3 |

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Table 1
(continued)

| Samples | $\mathrm{Th} / \mathrm{U}$ | ${ }^{207} \mathrm{~Pb} /{ }^{206} \mathrm{~Pb}$ | $2 \sigma$ | ${ }^{207} \mathrm{~Pb} /{ }^{235} \mathrm{U}$ | $2 \sigma$ | ${ }^{206} \mathrm{~Pb}{ }^{238} \mathrm{U}$ | $2 \sigma$ | ${ }^{207} \mathrm{~Pb} /{ }^{206} \mathrm{~Pb}$ | $2 \sigma$ | ${ }^{207} \mathrm{~Pb}{ }^{235} \mathrm{U}$ | $2 \sigma$ | ${ }^{206} \mathrm{~Pb}{ }^{238} \mathrm{U}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| D18-7-1-20 | 0.69 | 0.15785 | 0.00502 | 8.23613 | 0.32448 | 0.37117 | 0.00979 | 2433.0 | 49.1 | 2257.3 | 35.7 | 2034.9 |
| D18-6-1-14 | 0.72 | 0.16178 | 0.00561 | 6.04932 | 0.24938 | 0.26840 | 0.00642 | 2475.9 | 58.6 | 1983.0 | 35.9 | 1532.7 |
| D18-6-1-15 | 0.62 | 0.16914 | 0.00535 | 11.80579 | 0.37221 | 0.50268 | 0.00579 | 2549.1 | 52.8 | 2589.1 | 29.6 | 2625.3 |
| D18-6-1-16 | 0.90 | 0.16751 | 0.00494 | 11.50607 | 0.34429 | 0.49450 | 0.00569 | 2533.0 | 50.5 | 2565.1 | 28.0 | 2590.1 |
| D18-6-1-17 | 0.85 | 0.05660 | 0.01617 | 3.23449 | 0.67272 | 0.25154 | 0.00744 | 476.0 | 533.3 | 1465.5 | 161.3 | 1446.4 |
| D18-6-1-18 | 0.96 | 0.42498 | 0.01342 | 33.67022 | 4.55918 | 0.49549 | 0.05158 | 3999.8 | 46.4 | 3600.4 | 133.5 | 2594.4 |
| D18-6-1-19 | 0.60 | 0.16174 | 0.00447 | 9.30105 | 0.26023 | 0.41403 | 0.00487 | 2473.8 | 46.1 | 2368.1 | 25.7 | 2233.3 |
| D18-6-1-20 | 0.42 | 0.16276 | 0.00470 | 11.19395 | 0.32137 | 0.49472 | 0.00516 | 2484.3 | 48.8 | 2539.4 | 26.8 | 2591.1 |

TABLE 2
In-situ zircon Lu-Hf analytical results of the Biliya amphibolite (Samples D18-1-1, D18-2-1 and D18-7-1) and metadiorite

| Samples | Age (Ma) | ${ }^{176} \mathrm{Yb} /{ }^{177} \mathrm{Hf}$ | $2 \sigma$ | ${ }^{176} \mathrm{Lu} /{ }^{177} \mathrm{Hf}$ | $2 \sigma$ | ${ }^{176} \mathrm{Hf}{ }^{1 / 77} \mathrm{Hf}$ | $2 \sigma$ | ${ }^{176} \mathrm{Hf} /{ }^{177} \mathrm{Hf} \mathrm{i}_{\text {i }}$ | $\varepsilon_{\mathrm{Hf}}(0)$ | $\varepsilon_{\mathrm{Hf}}(\mathrm{t})$ | $\mathrm{T}_{\mathrm{DM}}(\mathrm{Ma})$ | $\mathrm{T}_{\mathrm{DM}}{ }^{\text {C }}$ (Ma) | $\mathrm{f}_{\text {LuHf }}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| D18-1-1-01 | 2481 | 0.02353 | 0.00032 | 0.00084 | 0.00001 | 0.28128 | 0.00002 | 0.28124 | -52.7 | 1.5 | 2738 | 2893 | -0.97 |
| D18-1-1-02 | 2481 | 0.02954 | 0.00025 | 0.00107 | 0.00001 | 0.28128 | 0.00002 | 0.28123 | -52.7 | 1.1 | 2754 | 2916 | -0.97 |
| D18-1-1-03 | 2481 | 0.02362 | 0.00047 | 0.00090 | 0.00001 | 0.28125 | 0.00002 | 0.28121 | -53.8 | 0.3 | 2787 | 2971 | -0.97 |
| D18-1-1-04 | 2481 | 0.02375 | 0.00037 | 0.00088 | 0.00001 | 0.28132 | 0.00002 | 0.28128 | -51.4 | 2.7 | 2692 | 2819 | -0.97 |
| D18-1-1-05 | 2481 | 0.03689 | 0.00112 | 0.00131 | 0.00004 | 0.28135 | 0.00002 | 0.28128 | -50.5 | 3.0 | 2685 | 2804 | -0.96 |
| D18-1-1-06 | 2481 | 0.02138 | 0.00061 | 0.00083 | 0.00003 | 0.28126 | 0.00002 | 0.28122 | -53.4 | 0.8 | 2764 | 2936 | -0.98 |
| D18-1-1-07 | 2481 | 0.02185 | 0.00078 | 0.00082 | 0.00003 | 0.28128 | 0.00002 | 0.28124 | -52.9 | 1.3 | 2746 | 2907 | -0.98 |
| D18-1-1-08 | 2481 | 0.03042 | 0.00107 | 0.00106 | 0.00003 | 0.28132 | 0.00002 | 0.28127 | -51.3 | 2.6 | 2699 | 2830 | -0.97 |
| D18-1-1-09 | 2481 | 0.02230 | 0.00059 | 0.00083 | 0.00003 | 0.28126 | 0.00002 | 0.28122 | -53.6 | 0.6 | 2772 | 2948 | -0.97 |
| D18-1-1-10 | 2481 | 0.02634 | 0.00069 | 0.00094 | 0.00002 | 0.28132 | 0.00002 | 0.28127 | -51.4 | 2.6 | 2697 | 2827 | -0.97 |
| D18-1-1-11 | 2481 | 0.03297 | 0.00106 | 0.00119 | 0.00002 | 0.28134 | 0.00002 | 0.28128 | -50.8 | 2.9 | 2689 | 2812 | -0.96 |
| D18-1-1-12 | 2481 | 0.02440 | 0.00058 | 0.00091 | 0.00003 | 0.28130 | 0.00002 | 0.28126 | -52.1 | 2.0 | 2720 | 2863 | -0.97 |
| D18-1-1-13 | 2481 | 0.02509 | 0.00028 | 0.00093 | 0.00001 | 0.28130 | 0.00002 | 0.28125 | -52.1 | 1.9 | 2723 | 2868 | -0.97 |
| D18-1-1-14 | 2481 | 0.02293 | 0.00018 | 0.00089 | 0.00001 | 0.28127 | 0.00002 | 0.28123 | -53.2 | 1.0 | 2760 | 2928 | -0.97 |
| D18-1-1-15 | 2481 | 0.02385 | 0.00028 | 0.00091 | 0.00001 | 0.28122 | 0.00002 | 0.28118 | -54.9 | -0.8 | 2827 | 3035 | -0.97 |
| D18-1-1-16 | 2481 | 0.02635 | 0.00020 | 0.00097 | 0.00001 | 0.28131 | 0.00002 | 0.28126 | -51.8 | 2.2 | 2712 | 2850 | -0.97 |
| D18-1-1-17 | 2481 | 0.06071 | 0.00360 | 0.00200 | 0.00011 | 0.28145 | 0.00003 | 0.28135 | -46.8 | 5.4 | 2592 | 2653 | -0.94 |
| D18-1-1-18 | 2481 | 0.02310 | 0.00052 | 0.00087 | 0.00002 | 0.28131 | 0.00002 | 0.28127 | -51.7 | 2.5 | 2702 | 2836 | -0.97 |
| D18-1-1-19 | 2481 | 0.03338 | 0.00060 | 0.00122 | 0.00003 | 0.28134 | 0.00002 | 0.28129 | -50.5 | 3.1 | 2682 | 2800 | -0.96 |
| D18-1-1-20 | 2481 | 0.02522 | 0.00040 | 0.00089 | 0.00001 | 0.28130 | 0.00002 | 0.28125 | -52.2 | 1.9 | 2722 | 2867 | -0.97 |
| D18-2-1-1 | 2539 | 0.01793 | 0.00060 | 0.00068 | 0.00002 | 0.28126 | 0.00002 | 0.28122 | -53.6 | 2.2 | 2762 | 2898 | -0.98 |
| D18-2-1-2 | 2539 | 0.02448 | 0.00020 | 0.00085 | 0.00000 | 0.28141 | 0.00002 | 0.28137 | -48.1 | 7.4 | 2564 | 2579 | -0.97 |
| D18-2-1-3 | 2539 | 0.03990 | 0.00271 | 0.00137 | 0.00007 | 0.28135 | 0.00002 | 0.28128 | -50.2 | 4.4 | 2681 | 2764 | -0.96 |
| D18-2-1-4 | 2539 | 0.02563 | 0.00028 | 0.00095 | 0.00001 | 0.28129 | 0.00002 | 0.28125 | -52.3 | 3.0 | 2732 | 2849 | -0.97 |
| D18-2-1-5 | 2539 | 0.03851 | 0.00420 | 0.00114 | 0.00010 | 0.28135 | 0.00003 | 0.28129 | -50.4 | 4.7 | 2669 | 2746 | -0.97 |
| D18-2-1-6 | 2539 | 0.06876 | 0.00122 | 0.00202 | 0.00003 | 0.28147 | 0.00002 | 0.28138 | -45.9 | 7.6 | 2557 | 2566 | -0.94 |
| D18-2-1-7 | 2539 | 0.02401 | 0.00071 | 0.00084 | 0.00002 | 0.28127 | 0.00002 | 0.28123 | -53.1 | 2.5 | 2752 | 2881 | -0.97 |
| D18-2-1-8 | 2539 | 0.03410 | 0.00055 | 0.00124 | 0.00002 | 0.28131 | 0.00002 | 0.28125 | -51.7 | 3.1 | 2728 | 2840 | -0.96 |
| D18-2-1-9 | 2539 | 0.03679 | 0.00210 | 0.00122 | 0.00006 | 0.28135 | 0.00002 | 0.28129 | -50.4 | 4.5 | 2676 | 2756 | -0.96 |
| D18-2-1-10 | 2539 | 0.03248 | 0.00039 | 0.00115 | 0.00002 | 0.28130 | 0.00002 | 0.28125 | -51.9 | 3.0 | 2731 | 2845 | -0.97 |
| D18-2-1-11 | 2539 | 0.03743 | 0.00097 | 0.00130 | 0.00005 | 0.28133 | 0.00003 | 0.28126 | -51.2 | 3.6 | 2712 | 2813 | -0.96 |
| D18-2-1-12 | 2539 | 0.01103 | 0.00005 | 0.00042 | 0.00000 | 0.28120 | 0.00002 | 0.28118 | -55.5 | 0.8 | 2813 | 2984 | -0.99 |
| D18-2-1-13 | 2539 | 0.02018 | 0.00018 | 0.00073 | 0.00002 | 0.28132 | 0.00003 | 0.28128 | -51.4 | 4.3 | 2682 | 2770 | -0.98 |

Table 2
(continued)

| Samples | Age (Ma) | ${ }^{176} \mathrm{Yb}{ }^{1 / 77} \mathrm{Hf}$ | $2 \sigma$ | ${ }^{176} \mathrm{Lu} /{ }^{177} \mathrm{Hf}$ | $2 \sigma$ | ${ }^{176} \mathrm{Hf} /{ }^{177} \mathrm{Hf}$ | $2 \sigma$ | ${ }^{176} \mathrm{Hf} /{ }^{177} \mathrm{Hf}{ }_{\text {i }}$ | $\varepsilon_{\text {Hff }}(0)$ | $\varepsilon_{\text {Hff }}(\mathrm{t})$ | $\mathrm{T}_{\mathrm{DM}}(\mathrm{Ma})$ | $\mathrm{T}_{\mathrm{DM}}{ }^{\text {c }}$ (Ma) | $\mathrm{f}_{\text {LuHf }}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| D18-2-1-14 | 2539 | 0.02344 | 0.00111 | 0.00079 | 0.00003 | 0.28133 | 0.00002 | 0.28129 | -51.1 | 4.5 | 2672 | 2753 | -0.98 |
| D18-2-1-15 | 2539 | 0.02968 | 0.00049 | 0.00106 | 0.00001 | 0.28129 | 0.00002 | 0.28124 | -52.4 | 2.7 | 2745 | 2867 | -0.97 |
| D18-2-1-16 | 2539 | 0.01641 | 0.00021 | 0.00056 | 0.00000 | 0.28126 | 0.00002 | 0.28124 | -53.3 | 2.6 | 2743 | 2870 | -0.98 |
| D18-2-1-17 | 2539 | 0.03151 | 0.00077 | 0.00114 | 0.00003 | 0.28126 | 0.00002 | 0.28120 | -53.5 | 1.5 | 2792 | 2942 | -0.97 |
| D18-2-1-18 | 2539 | 0.05117 | 0.00270 | 0.00161 | 0.00006 | 0.28139 | 0.00003 | 0.28131 | -49.0 | 5.2 | 2648 | 2711 | -0.95 |
| D18-2-1-19 | 2539 | 0.02384 | 0.00066 | 0.00088 | 0.00003 | 0.28127 | 0.00002 | 0.28123 | -53.2 | 2.3 | 2759 | 2892 | -0.97 |
| D18-2-1-20 | 2539 | 0.03100 | 0.00048 | 0.00103 | 0.00001 | 0.28130 | 0.00002 | 0.28125 | -52.2 | 3.0 | 2732 | 2847 | -0.97 |
| D18-7-1-1 | 2539 | 0.02284 | 0.00022 | 0.00083 | 0.00001 | 0.28136 | 0.00002 | 0.28132 | -50.0 | 5.5 | 2634 | 2692 | -0.98 |
| D18-7-1-2 | 2539 | 0.02103 | 0.00083 | 0.00073 | 0.00003 | 0.28128 | 0.00002 | 0.28125 | -52.7 | 3.0 | 2730 | 2847 | -0.98 |
| D18-7-1-3 | 2539 | 0.02425 | 0.00064 | 0.00100 | 0.00004 | 0.28129 | 0.00002 | 0.28125 | -52.2 | 3.0 | 2732 | 2848 | -0.97 |
| D18-7-1-4 | 2539 | 0.02790 | 0.00036 | 0.00109 | 0.00002 | 0.28126 | 0.00003 | 0.28120 | -53.5 | 1.5 | 2789 | 2938 | -0.97 |
| D18-7-1-5 | 2539 | 0.03070 | 0.00036 | 0.00112 | 0.00002 | 0.28137 | 0.00002 | 0.28132 | -49.6 | 5.4 | 2639 | 2698 | -0.97 |
| D18-7-1-6 | 2539 | 0.05111 | 0.00191 | 0.00166 | 0.00005 | 0.28142 | 0.00002 | 0.28134 | -47.9 | 6.2 | 2610 | 2651 | -0.95 |
| D18-7-1-7 | 2539 | 0.02716 | 0.00080 | 0.00101 | 0.00004 | 0.28133 | 0.00003 | 0.28128 | -51.1 | 4.1 | 2690 | 2780 | -0.97 |
| D18-7-1-8 | 2539 | 0.02561 | 0.00070 | 0.00095 | 0.00002 | 0.28130 | 0.00002 | 0.28125 | -52.1 | 3.2 | 2724 | 2834 | -0.97 |
| D18-7-1-9 | 2539 | 0.03559 | 0.00080 | 0.00135 | 0.00004 | 0.28132 | 0.00003 | 0.28126 | -51.3 | 3.3 | 2721 | 2827 | -0.96 |
| D18-7-1-10 | 2539 | 0.03819 | 0.00086 | 0.00146 | 0.00005 | 0.28136 | 0.00004 | 0.28129 | -50.0 | 4.4 | 2680 | 2761 | -0.96 |
| D18-3-1-11 | 2537 | 0.01213 | 0.00074 | 0.00048 | 0.00003 | 0.28123 | 0.00002 | 0.28121 | -54.5 | 1.6 | 2781 | 2933 | -0.99 |
| D18-3-1-12 | 2537 | 0.02609 | 0.00051 | 0.00096 | 0.00001 | 0.28130 | 0.00002 | 0.28126 | -51.9 | 3.4 | 2717 | 2825 | -0.97 |
| D18-3-1-13 | 2537 | 0.02412 | 0.00015 | 0.00081 | 0.00002 | 0.28128 | 0.00002 | 0.28124 | -52.7 | 2.9 | 2735 | 2855 | -0.98 |
| D18-3-1-14 | 2537 | 0.02651 | 0.00069 | 0.00097 | 0.00002 | 0.28123 | 0.00002 | 0.28118 | -54.6 | 0.6 | 2822 | 2993 | -0.97 |
| D18-3-1-15 | 2537 | 0.03260 | 0.00056 | 0.00106 | 0.00002 | 0.28131 | 0.00002 | 0.28126 | -51.8 | 3.3 | 2720 | 2829 | -0.97 |
| D18-3-1-16 | 2537 | 0.03124 | 0.00036 | 0.00111 | 0.00001 | 0.28123 | 0.00002 | 0.28117 | -54.7 | 0.3 | 2834 | 3011 | -0.97 |
| D18-3-1-17 | 2537 | 0.02836 | 0.00153 | 0.00094 | 0.00004 | 0.28128 | 0.00002 | 0.28123 | -52.9 | 2.4 | 2754 | 2885 | -0.97 |
| D18-3-1-18 | 2537 | 0.04164 | 0.00065 | 0.00132 | 0.00003 | 0.28136 | 0.00002 | 0.28129 | -50.0 | 4.6 | 2668 | 2745 | -0.96 |
| D18-3-1-19 | 2537 | 0.03852 | 0.00038 | 0.00148 | 0.00002 | 0.28140 | 0.00004 | 0.28133 | -48.6 | 5.8 | 2625 | 2675 | -0.96 |
| D18-3-1-20 | 2537 | 0.02953 | 0.00099 | 0.00109 | 0.00002 | 0.28127 | 0.00002 | 0.28122 | -53.0 | 2.1 | 2766 | 2903 | -0.97 |
| D18-6-1-1 | 2565 | 0.01292 | 0.00016 | 0.00050 | 0.00000 | 0.28119 | 0.00002 | 0.28117 | -55.9 | 0.7 | 2837 | 3005 | -0.98 |
| D18-6-1-2 | 2565 | 0.01516 | 0.00013 | 0.00060 | 0.00000 | 0.28123 | 0.00002 | 0.28120 | -54.4 | 2.1 | 2787 | 2923 | -0.98 |
| D18-6-1-3 | 2565 | 0.02363 | 0.00014 | 0.00089 | 0.00000 | 0.28127 | 0.00002 | 0.28123 | -53.1 | 2.9 | 2757 | 2874 | -0.97 |
| D18-6-1-4 | 2565 | 0.02408 | 0.00017 | 0.00084 | 0.00001 | 0.28125 | 0.00002 | 0.28121 | -53.7 | 2.4 | 2778 | 2907 | -0.97 |
| D18-6-1-5 | 2565 | 0.02067 | 0.00038 | 0.00069 | 0.00001 | 0.28125 | 0.00002 | 0.28121 | -53.9 | 2.5 | 2773 | 2901 | -0.98 |
| D18-6-1-6 | 2565 | 0.01332 | 0.00032 | 0.00051 | 0.00001 | 0.28125 | 0.00002 | 0.28123 | -53.8 | 2.9 | 2757 | 2876 | -0.98 |
| D18-6-1-7 | 2565 | 0.01391 | 0.00052 | 0.00054 | 0.00002 | 0.28123 | 0.00002 | 0.28120 | -54.6 | 2.1 | 2788 | 2926 | -0.98 |
| D18-6-1-8 | 2565 | 0.01710 | 0.00021 | 0.00065 | 0.00000 | 0.28126 | 0.00002 | 0.28123 | -53.5 | 2.9 | 2755 | 2872 | -0.98 |

TABLE 2
(continued)

| Samples | Age (Ma) | ${ }^{176} \mathrm{Yb} /{ }^{177} \mathrm{Hf}$ | $2 \sigma$ | ${ }^{176} \mathrm{Lu} /{ }^{177} \mathrm{Hf}$ | $2 \sigma$ | ${ }^{176} \mathrm{Hf} /{ }^{177} \mathrm{Hf}$ | $2 \sigma$ | ${ }^{176} \mathrm{Hf} /{ }^{177} \mathrm{Hf}_{\mathrm{i}}$ | $\varepsilon_{\mathrm{Hf}}(0)$ | $\varepsilon_{\mathrm{Hff}}(\mathrm{t})$ | $\mathrm{T}_{\mathrm{DM}}(\mathrm{Ma})$ | $\mathrm{T}_{\mathrm{DM}}(\mathrm{Ma})$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| D18-6-1-9 | 2565 | 0.02344 | 0.00035 | 0.00087 | 0.00001 | 0.28125 | 0.00002 | 0.28121 | -53.9 | 2.2 | 2786 | 2919 |
| $\mathrm{f}_{\mathrm{L} / \mathrm{Hff}}$ |  |  |  |  |  |  |  |  |  |  |  |  |
| D18-6-1-10 | 2565 | 0.02413 | 0.00016 | 0.00093 | 0.00001 | 0.28128 | 0.00002 | 0.28123 | -52.8 | 3.2 | 2748 | 2858 |
| D18-6-1-11 | 2565 | 0.01946 | 0.00051 | 0.00077 | 0.00002 | 0.28125 | 0.00003 | 0.28121 | -53.8 | 2.4 | 2774 | 2902 |
| D18-6-1-12 | 2565 | 0.02870 | 0.00056 | 0.00096 | 0.00001 | 0.28127 | 0.00002 | 0.28123 | -53.0 | 2.9 | 2758 | 2874 |
| D18-6-1-13 | 2565 | 0.01769 | 0.00024 | 0.00068 | 0.00001 | 0.28127 | 0.00002 | 0.28123 | -53.3 | 3.1 | 2749 | 2862 |
| D18-6-1-14 | 2565 | 0.01794 | 0.00140 | 0.00065 | 0.00005 | 0.28131 | 0.00002 | 0.28127 | -51.8 | 4.6 | 2692 | 2771 |
| D18-6-1-15 | 2565 | 0.02313 | 0.00039 | 0.00086 | 0.00001 | 0.28129 | 0.00002 | 0.28125 | -52.4 | 3.6 | 2729 | 2828 |
| D18-6-1-16 | 2565 | 0.01985 | 0.00015 | 0.00075 | 0.00001 | 0.28126 | 0.00002 | 0.28122 | -53.5 | 2.7 | 2764 | 2885 |
| D18-6-1-17 | 2565 | 0.02915 | 0.00019 | 0.00109 | 0.00001 | 0.28133 | 0.00002 | 0.28128 | -50.8 | 4.8 | 2685 | 2756 |
| D18-6-1-18 | 2565 | 0.01621 | 0.00017 | 0.00061 | 0.00001 | 0.28128 | 0.00002 | 0.28125 | -52.9 | 3.6 | 2730 | 2832 |
| D18-6-1-19 | 2565 | 0.03925 | 0.00099 | 0.00129 | 0.00003 | 0.28142 | 0.00003 | 0.28136 | -47.7 | 7.7 | 2575 | 2581 |
| D18-6-1-20 | 2565 | 0.02162 | 0.00020 | 0.00079 | 0.00001 | 0.28124 | 0.00002 | 0.28120 | -54.1 | 2.0 | 2790 | 2927 |



Fig. 4. (A) SHRIMP zircon U-Pb Concordia diagram. (B-E) LA-ICP-MS zircon U-Pb Concordia diagrams. (F) Th and U contents of the analyzed zircons.
igneous origin (Hanchar and Miller, 1993). The magmatic cores yielded an upper concordia intercept age of $2539 \pm 23 \mathrm{Ma}(\mathrm{MSWD}=3.4, \mathrm{n}=20$ ) and a weighted mean $207 \mathrm{~Pb} / 206 \mathrm{~Pb}$ age of $2529 \pm 18 \mathrm{Ma}(\mathrm{MSWD}=0.87, \mathrm{n}=6$; fig. 4B). This age ( $2539 \pm$ 23 Ma ) represents the magma crystallization age of the protolith. Twenty in situ Lu-Hf analyses were recalculated to this crystallization age. They show initial $176 \mathrm{Hf} / 177 \mathrm{Hf}$ ratios of 0.28118 to 0.28138 and $\varepsilon H f(t)$ values of -0.1 to +6.7 (fig. 5 A ). The twostaged model ages are in the range of 2588 Ma to 3009 Ma .

Sample D18-7-1: Twenty analyses have been carried out on 20 zircon grains from this metadiorite sample. The twenty zircon crystals are euhedral to subhedral, with length between 60 and $110 \mu \mathrm{~m}$, and contain broad cores and oscillatory thinner rims (fig. 4C). The bright metamorphic rims are too narrow to be analyzed by LA-ICPMS. Our analyses concentrate at the cores and show high $\mathrm{Th} / \mathrm{U}$ ratios of 0.5 to 0.9 (fig. 4G). These characteristics indicate an igneous origin (Hanchar and Miller, 1993). The magmatic cores define an upper intercept age of $2539 \pm 25 \mathrm{Ma}$ (MSWD = 1.2, n $=20$ ) and a weighted mean $207 \mathrm{~Pb} / 206 \mathrm{~Pb}$ age of $2527 \pm 35 \mathrm{Ma}(\mathrm{MSWD}=1.5, \mathrm{n}=10$; fig. 4C), with the former age ( $2539 \pm 25 \mathrm{Ma}$ ) taken as the magma crystallization age of the granodiorite. Twenty in situ Lu -Hf analyses were recalculated to this crystallization


Fig. 5. U-Pb ages (Ma) vs. zircon $\varepsilon H f(t)$ values (A) and whole-rock $\varepsilon N d(t)$ vs. whole-rock $\varepsilon H f(t)$ values (B). Shaded areas in (A) are from Yang and others (2008). The mantle and crustal arrays are from Albarède and others (2000) and Vervoort and others (2000). The MORB and OIB fields are from Georoc (http://georoc.mpch-mainz.gwdg.de/georoc/).
age. They show initial $176 \mathrm{Hf} / 177 \mathrm{Hf}$ ratios of 0.28117 to 0.28134 and $\varepsilon \mathrm{Hf}(\mathrm{t})$ values of 0.6 to +5.4 (fig. 5 A ). The two-staged model ages are in the range of 2673 Ma to 3037 Ma .

Sample D18-3-1: Twenty analyses were performed on 20 zircon grains. The 20 spots display well-preserved euhedral growth zones and oscillatory zoning, with length between 100 and $180 \mu \mathrm{~m}$ and aspect ratios of 1:1 to 2:1 (fig. 4D). They show high Th/U ratios of 0.4 to 1.2 (fig. 4G and table 1), indicative of a magmatic origin (Hanchar and Miller, 1993). Analyses of these zircons gave an upper intercept age of $2537 \pm 17 \mathrm{Ma}$ (MSWD = $1.4, \mathrm{n}=20$ ), and a weighted mean $207 \mathrm{~Pb} / 206 \mathrm{~Pb}$ age of $2510 \pm 18 \mathrm{Ma}($ MSWD $=0.7, \mathrm{n}=$ 7; fig. 4D). This age ( $2537 \pm 17 \mathrm{Ma}$ ) represents the magma crystallization age of the protolith. Twenty in situ Lu-Hf analyses were recalculated to this crystallization age. They show initial $176 \mathrm{Hf} / 177 \mathrm{Hf}$ ratios of 0.28117 to 0.28133 and $\varepsilon \mathrm{Hf}(\mathrm{t})$ values of -0.5 to +5.0 (fig. 5A). The two-staged model ages are in the range of 2697 Ma to 3033 Ma .

Sample D18-6-1: Twenty analyses were performed on 20 zircon grains. The 20 spots display well-preserved euhedral growth zones and oscillatory zoning, with length between 80 and $150 \mu \mathrm{~m}$ and aspect ratios of 1:1 to 3:1 (fig. 4E). They show high $\mathrm{Th} / \mathrm{U}$ ratios of 0.4 to 1.0 (fig. 4G and table 1), which indicates a magmatic origin (Hanchar and Miller, 1993). They gave an upper intercept age of $2565 \pm 26 \mathrm{Ma}$ (MSWD = 1.2, $\mathrm{n}=20$ ), and a weighted mean $207 \mathrm{~Pb} / 206 \mathrm{~Pb}$ age of $2549 \pm 32 \mathrm{Ma}(\mathrm{MSWD}=1.3, \mathrm{n}=9$; fig. 4 E ), and the former age is taken as the crystallization age of the protolith. Twenty in situ Lu-Hf analyses were recalculated to this crystallization age. They show initial
$176 \mathrm{Hf} / 177 \mathrm{Hf}$ ratios of 0.28117 to 0.28136 and $\varepsilon \mathrm{Hf}(\mathrm{t})$ values of -0.7 to +6.2 (fig. 5 A ). The two-staged model ages are in the range of 2619 Ma to 3046 Ma .

## Geochemical Analytical Results

Nine amphibolite and six metadiorite samples were analyzed for their major and trace element compositions, and the results are listed in table 3. The amphibolite samples have $\mathrm{SiO} 2=44.24-51.12 \mathrm{wt} . \%, \mathrm{Al2O} 3=12.39-14.53 \mathrm{wt} . \%, \mathrm{Fe} 2 \mathrm{O} 3 \mathrm{~T}=11.89-13.43$ wt. $\%, \mathrm{CaO}=11.31-14.17 \mathrm{wt} . \%, \mathrm{MgO}=4.83-5.95 \mathrm{wt} . \%, \mathrm{Na} 2 \mathrm{O}=1.79-3.08 \mathrm{wt} . \%$, $\mathrm{Na} 2 \mathrm{O} / \mathrm{K} 2 \mathrm{O}=10-18$. In the SiO 2 vs. $\mathrm{Nb} / \mathrm{Y}$ diagram, the samples plot in the gabbro field (fig. 6A). These samples define a typical tholeiitic trend in the SiO 2 vs. $\mathrm{FeOT} /$ MgO diagram (fig. 6B). The amphibolite samples are depleted in highly incompatible elements and show left-leaning smooth primitive mantle-normalized multi-element patterns, and weakly left-leaning chondrite-normalized REE patterns $((\mathrm{La} / \mathrm{Sm}) \mathrm{CN}=$ $0.92-1.07$, $(\mathrm{La} / \mathrm{Yb}) \mathrm{CN}=1.07-1.20)$ and negative Eu anomalies $\left(\mathrm{Eu}^{*}=0.62-0.77\right.$; fig. 7). The bulk rock samples have $\varepsilon N d(t)=-1.7-+4.5$ and $\varepsilon H f(t)=-1.9$ (fig. 5 B and table 4).

The six metadiorite samples have $\mathrm{SiO} 2=59.07-64.56 \mathrm{wt} . \%, \mathrm{MgO}=2.27-4.12 \mathrm{wt}$. $\%$, and $(\mathrm{K} 2 \mathrm{O}+\mathrm{Na} 2 \mathrm{O})=5.31-7.16 \mathrm{wt} . \%$. In the QAP and An-Ab-Or diagrams, all six samples fall into the granodiorite field. They are weakly peraluminous $(\mathrm{A} / \mathrm{CNK}=$ $0.84-1.18)$ and high-K calc-alkaline. The metadiorite samples have strongly negative anomalies of $\mathrm{Nb}\left(\mathrm{Nb}^{*}=0.15-0.31\right)$, Ta and $\mathrm{Sr}\left(\mathrm{Sr}^{*}=0.56-0.73\right.$; fig. 7 A$)$, and moderately negative to positive Eu anomalies ( $\mathrm{Eu}^{*}=0.77-1.20$ ) with right-leaning REE patterns (fig. 7B). The metadiorite samples analyzed for their whole-rock Sm-Nd and LuHf isotopes yielded $\varepsilon \mathrm{Nd}(\mathrm{t})=-0.5-+3.6$ and TDM2 age $=2569-2910 \mathrm{Ma}, \varepsilon \mathrm{Hf}(\mathrm{t})=$ $+0.5-+1.4$ and TDM2 age $=2916-2968 \mathrm{Ma}$ (fig. 5B and table 4).

## DISCUSSION

## Neoarchean Basement in the Microcontinents within CAOB

Recently, late Paleoproterozoic ( 1.8 Ga ) rocks were identified in the CAOB (Wang and others, 2006; Pei and others, 2007; Sun and others, 2013), for example, $\sim 1.8$ Ga granites in the borehole of the Songliao Basin (Wang and others, 2006; Pei and others, 2007), $\sim 1.84$ to 1.7 Ga granitic gneisses in the Huma area of the Xing'an Terrane (Sun and others, 2013), $\sim 1.7$ Ga diorite porphyrite and diorite in the central Mongolia Terrane, and $\sim 1.85$ to 1.82 Ga crystalline rocks in the Dzabkhan Terrane in Russia (Kozakov and others, 2007, 2011). Early Mesoproterozoic (1.4 Ga) rocks are also found throughout the CAOB, such as the 1.45 to 1.40 Ga granitoids in the Jiujing region, 1.45 to 1.40 Ga granitoids and amphibolites from the Xingxingxia, Weiya and Alatage areas in the Central Tianshan, 1.46 to 1.43 Ga supracrustal rocks and intrusive rocks in northern Alxa Terrane, $\sim 1.37$ Ga rhyolites at Kyrgyz of North Tianshan, and 1.40 to 1.36 Ga granitoids in Xilinhot Terrane (for example, Kröner and others, 2013; Han and others, 2017; Yuan and others, 2019). In addition, Liu and others (2020) reported one early Neoproterozoic granitic pluton (964-947 Ma) in the Erguna Terrane.

In contrast, confirmed Archean rocks are scarce in the eastern CAOB (Zhou and others, 2018 and references therein). This study has identified the coexistence of Neoarchean amphibolite and metadiorite in the Erguna Terrane (eastern CAOB). Similarly, in the Junggar Terrane (western CAOB), basement enclaves in Ordovician volcanic rocks yielded a Neoarchean magmatic zircon U-Pb age ( 2522 Ma ; Xu and others, 2015). In the central Tianshan Terrane (western CAOB), the protolith age of the orthogneiss has been dated to be 2529 to 2513 Ma by Wang and others (2017). In the central CAOB, a dioritic gneiss yielded magmatic ages of 2546 to 2520 Ma (Kröner and others, 2015). From the detrital zircon age record, Paleozoic sequences

Table 3
Major and trace element data of the Biliya amphibolite and metadiorite

| Samples | Metadiorite |  |  |  |  |  | Amphibolite |  |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | D18- | D18- | D18- | D18- | D18- |  |  |  |  |  | D18- | D18- | D18- |  | D18 |
|  | 1-1 | 1-2 | 2-1 | 2-2 | 7-1 | 8-1 | 3-1 | 3-2 | 4-1 | 4-2 | 4-3 | 5-1 | 5-2 | 6-1 | -6-2 |
| $\mathrm{SiO}_{2}$ | 64.14 | 64.13 | 64.56 | 64.36 | 59.53 | 59.07 | 45.77 | 44.81 | 47.12 | 47.25 | 46.43 | 44.96 | 44.24 | 51.12 | 50.79 |
| $\mathrm{TiO}_{2}$ | 0.53 | 0.56 | 0.52 | 0.55 | 1.13 | 1.17 | 0.97 | 1.00 | 1.11 | 1.11 | 1.16 | 1.09 | 1.13 | 0.96 | 1.00 |
| $\mathrm{Al}_{2} \mathrm{O}_{3}$ | 16.31 | 16.23 | 14.82 | 14.53 | 15.39 | 15.19 | 12.78 | 12.39 | 14.49 | 14.53 | 14.26 | 14.12 | 13.84 | 12.82 | 12.54 |
| $\mathrm{Fe}_{2} \mathrm{O}_{3 \mathrm{~T}}$ | 5.91 | 6.03 | 6.85 | 7.11 | 7.54 | 7.85 | 12.74 | 12.97 | 13.07 | 13.09 | 13.35 | 13.01 | 13.43 | 11.89 | 12.21 |
| MnO | 0.08 | 0.09 | 0.13 | 0.14 | 0.12 | 0.13 | 0.21 | 0.21 | 0.19 | 0.19 | 0.20 | 0.20 | 0.21 | 0.18 | 0.18 |
| MgO | 2.44 | 2.59 | 2.27 | 2.53 | 3.88 | 4.12 | 5.44 | 5.69 | 5.37 | 5.39 | 5.73 | 5.57 | 5.95 | 4.83 | 5.05 |
| CaO | 2.40 | 2.44 | 3.51 | 3.71 | 5.42 | 5.68 | 13.60 | 14.17 | 11.62 | 11.61 | 12.12 | 12.88 | 13.45 | 11.31 | 11.76 |
| $\mathrm{Na}_{2} \mathrm{O}$ | 3.62 | 3.69 | 3.20 | 3.26 | 2.96 | 2.98 | 2.15 | 2.11 | 2.78 | 2.77 | 2.81 | 2.59 | 2.58 | 2.40 | 2.38 |
| $\mathrm{K}_{2} \mathrm{O}$ | 3.19 | 3.48 | 2.11 | 2.32 | 2.49 | 2.74 | 0.21 | 0.21 | 0.17 | 0.17 | 0.16 | 0.21 | 0.20 | 0.16 | 0.15 |
| $\mathrm{P}_{2} \mathrm{O}_{5}$ | 0.17 | 0.17 | 0.17 | 0.17 | 0.37 | 0.37 | 0.08 | 0.08 | 0.09 | 0.09 | 0.09 | 0.09 | 0.08 | 0.08 | 0.07 |
| LOI | 0.91 | 1.30 | 1.50 | 1.96 | 0.86 | 1.49 | 5.70 | 6.92 | 3.60 | 3.59 | 4.72 | 4.84 | 6.04 | 3.88 | 4.96 |
| SUM | 99.71 | 100.19 | 99.62 | 100.06 | 99.69 | 100.19 | 99.67 | 99.48 | 99.60 | 99.79 | 99.84 | 99.56 | 99.99 | 99.62 | 100.04 |
| Li | 31.0 | 31.2 | 18.7 | 17.4 | 39.9 | 38.8 | 14.8 | 14.0 | 13.0 | 12.7 | 11.8 | 13.8 | 12.3 | 11.7 | 10.5 |
| Sc | 12.3 | 12.2 | 15.2 | 14.5 | 17.6 | 17.0 | 39.6 | 48.2 | 44.0 | 42.6 | 51.0 | 45.0 | 52.9 | 39.1 | 46.2 |
| V | 82.8 | 70.8 | 94.3 | 81.3 | 118 | 107 | 305 | 259 | 349 | 330 | 288 | 338 | 276 | 301 | 256 |
| Cr | 71.8 | 59.7 | 79.1 | 75.5 | 106 | 89.5 | 144 | 173 | 164 | 157 | 18 | 161 | 176 | 140 | 165 |
| Co | 16.3 | 15.4 | 19.5 | 18.0 | 18.9 | 23.0 | 48.3 | 43.5 | 52.7 | 51.0 | 51.6 | 52.5 | 65.5 | 47.9 | 44.9 |
| Ni | 41.2 | 36.3 | 54.0 | 44.6 | 26.9 | 25.8 | 109 | 96.3 | 121 | 115 | 104 | 112 | 98.0 | 109 | 90.8 |
| Ga | 20.2 | 21.0 | 17.4 | 18.7 | 19.7 | 21.0 | 15.7 | 17.5 | 17.0 | 16.2 | 19.4 | 16.5 | 18.9 | 15.1 | 17.5 |
| Rb | 118 | 121 | 84.5 | 86.7 | 104 | 106 | 5.02 | 6.00 | 3.50 | 3.36 | 5.11 | 5.52 | 7.50 | 3.58 | 5.67 |
| Sr | 382 | 380 | 290 | 285 | 566 | 550 | 117 | 112 | 157 | 151 | 143 | 136 | 134 | 123 | 123 |
| Y | 15.8 | 16.7 | 22.0 | 21.9 | 26.3 | 27.5 | 22.9 | 24.6 | 24.0 | 23.4 | 25.3 | 24.2 | 25.5 | 23.3 | 24.1 |
| Zr | 113 | 110 | 137 | 121 | 262 | 236 | 54.4 | 69.1 | 60.7 | 59.5 | 74.5 | 58.9 | 73.7 | 51.2 | 62.0 |
| Nb | 7.96 | 8.36 | 6.22 | 7.75 | 18.1 | 18.8 | 2.94 | 3.34 | 3.04 | 3.13 | 2.89 | 3.12 | 3.21 | 2.87 | 2.76 |
| Cs | 6.75 | 10.2 | 3.83 | 5.43 | 5.91 | 9.05 | 0.097 | 0.22 | 0.037 | 0.047 | 0.10 | 0.067 | 0.12 | 0.043 | 0.094 |
| Ba | 849 | 936 | 437 | 522 | 786 | 850 | 67.0 | 96.5 | 68.1 | 67.2 | 87.8 | 61.7 | 105 | 57.3 | 87.7 |
| La | 29.1 | 32.6 | 24.4 | 26.8 | 38.4 | 43.4 | 3.68 | 4.28 | 3.86 | 3.54 | 4.12 | 3.82 | 4.11 | 3.69 | 4.12 |
| Ce | 57.9 | 63.5 | 49.9 | 53.8 | 82.2 | 91.5 | 9.53 | 10.8 | 10.3 | 9.69 | 10.9 | 10.0 | 11.0 | 9.82 | 10.5 |
| Pr | 6.69 | 7.52 | 5.63 | 6.47 | 9.57 | 11.2 | 1.51 | 1.74 | 1.58 | 1.49 | 1.78 | 1.56 | 1.84 | 1.46 | 1.71 |
| Nd | 25.9 | 28.2 | 21.2 | 24.1 | 37.1 | 41.0 | 7.46 | 8.55 | 7.56 | 7.85 | 8.79 | 8.16 | 9.16 | 7.41 | 8.57 |
| Sm | 4.26 | 4.60 | 4.12 | 4.05 | 6.78 | 6.92 | 2.38 | 2.58 | 2.48 | 2.35 | 2.54 | 2.57 | 2.79 | 2.59 | 2.48 |
| Eu | 1.19 | 1.43 | 0.95 | 1.36 | 1.89 | 2.57 | 0.71 | 0.71 | 0.67 | 0.68 | 0.64 | 0.71 | 0.64 | 0.59 | 0.60 |
| Gd | 3.61 | 3.97 | 3.49 | 3.65 | 5.82 | 6.25 | 3.28 | 3.55 | 3.56 | 3.41 | 3.55 | 3.55 | 3.57 | 3.54 | 3.49 |
| Tb | 0.50 | 0.57 | 0.54 | 0.57 | 0.81 | 0.87 | 0.59 | 0.58 | 0.61 | 0.59 | 0.58 | 0.64 | 0.59 | 0.60 | 0.55 |
| Dy | 2.83 | 2.90 | 3.50 | 3.41 | 4.64 | 4.63 | 3.71 | 3.84 | 3.91 | 3.95 | 4.07 | 4.16 | 4.06 | 3.87 | 3.83 |
| Ho | 0.54 | 0.55 | 0.75 | 0.71 | 0.96 | 0.92 | 0.85 | 0.83 | 0.88 | 0.85 | 0.86 | 0.90 | 0.86 | 0.85 | 0.80 |
| Er | 1.51 | 1.50 | 2.29 | 2.02 | 2.66 | 2.44 | 2.40 | 2.28 | 2.72 | 2.43 | 2.42 | 2.50 | 2.37 | 2.50 | 2.18 |
| Tm | 0.24 | 0.27 | 0.36 | 0.38 | 0.40 | 0.45 | 0.37 | 0.43 | 0.43 | 0.38 | 0.45 | 0.42 | 0.45 | 0.39 | 0.42 |
| Yb | 1.47 | 1.68 | 2.42 | 2.42 | 2.48 | 2.70 | 2.45 | 2.62 | 2.49 | 2.36 | 2.76 | 2.35 | 2.51 | 2.20 | 2.52 |
| Lu | 0.24 | 0.21 | 0.36 | 0.34 | 0.38 | 0.33 | 0.36 | 0.37 | 0.36 | 0.37 | 0.36 | 0.37 | 0.37 | 0.36 | 0.33 |
| Hf | 3.27 | 3.23 | 3.81 | 3.20 | 6.02 | 6.29 | 1.59 | 1.62 | 1.82 | 1.77 | 1.60 | 1.78 | 1.57 | 1.43 | 1.49 |
| Ta | 0.61 | 0.65 | 0.47 | 0.61 | 1.15 | 1.19 | 0.20 | 0.20 | 0.20 | 0.19 | 0.23 | 0.22 | 0.22 | 0.19 | 0.21 |
| Th | 7.71 | 7.27 | 6.90 | 6.30 | 10.5 | 9.19 | 0.33 | 0.39 | 0.34 | 0.29 | 0.38 | 0.29 | 0.32 | 0.26 | 0.27 |
| U | 2.25 |  | 2.09 |  | 1.91 |  | 0.094 |  | 0.084 | 0.071 |  | 0.074 |  | 0.074 |  |
| Mg\# | 49 | 50 | 44 | 45 | 55 | 55 | 50 | 51 | 49 | 49 | 50 | 50 | 51 | 49 | 49 |
| A/CNK | 1.18 | 1.14 | 1.06 | 0.99 | 0.88 | 0.84 | 0.45 | 0.42 | 0.56 | 0.56 | 0.53 | 0.51 | 0.48 | 0.52 | 0.49 |
| Eu* | 0.93 | 1.02 | 0.77 | 1.08 | 0.92 | 1.20 | 0.77 | 0.72 | 0.68 | 0.73 | 0.65 | 0.72 | 0.62 | 0.60 | 0.62 |
| Q | 24.31 | 22.22 | 30.08 | 27.58 | 18.71 | 16.67 | 5.63 | 3.79 | 5.14 | 5.23 | 3.02 | 1.77 | 0.00 | 14.60 | 13.58 |
| Or | 20.28 | 22.01 | 13.65 | 15.00 | 16.15 | 17.69 | 1.54 | 1.52 | 1.19 | 1.22 | 1.12 | 1.54 | 1.45 | 1.12 | 1.04 |
| Ab | 32.99 | 33.42 | 29.67 | 30.10 | 27.47 | 27.56 | 22.42 | 22.13 | 28.32 | 28.25 | 28.64 | 26.81 | 26.69 | 24.18 | 23.96 |
| An | 11.60 | 11.78 | 17.87 | 18.93 | 23.36 | 21.86 | 30.28 | 29.42 | 32.04 | 32.11 | 31.17 | 32.17 | 31.36 | 28.31 | 27.53 |

$\mathrm{Mg}^{\#}=\mathrm{MgO} / 40.3 /(\mathrm{MgO} / 40.3+\mathrm{FeOt} / 71.9), \mathrm{A} / \mathrm{CNK}=\mathrm{Al}_{2} \mathrm{O}_{3} / 101.9 /\left(\mathrm{CaO} / 56.1+\mathrm{Na}_{2} \mathrm{O} / 62.0+\mathrm{K}_{2} \mathrm{O} / 94.2\right)$


Fig. 6. (A) $\mathrm{SiO} 2-\mathrm{Nb} / \mathrm{Y}$ diagram and (C) AFM plot $(\mathrm{A}=\mathrm{Na} 2 \mathrm{O}+\mathrm{K} 2 \mathrm{O} ; \mathrm{F}=\mathrm{FeO}+0.8998 * \mathrm{Fe} 2 \mathrm{O} 3 ; \mathrm{M}=\mathrm{MgO})$.
in the Songliao Terrane (eastern CAOB) yielded a major age peak at 2585 Ma (Zhou and others, 2012). Zhou and others (2018) summarized 390 age data from the Paleozoic sequences in the Kazakhstan-Yili-Central Tianshan Terrane (western CAOB) and identified an age peak at 2510 Ma . In the Beishan Terrane (western CAOB), a slightly younger 2479 Ma age peak was also identified (Zhou and others, 2018). Xu and others (2017) and Liu and others (2020) also reported some detrital zircons with Neoarchean Hf model-ages (2598-2457 Ma). The Paleo-Asian Ocean is widely believed to have closed in late Permian to Triassic (Zhou and Wilde, 2013; Xiao and others, 2015), and these late Archean detrital materials in these terranes with the CAOB were most likely local-sourced. Thus, these Neoarchean igneous and detrital zircons support the contention that Archean basement is present in the microcontinents within the CAOB.

## Petrogenesis of the MORB-like Amphibolite: ~ 20 \% Partial Melting of Lithospheric Mantle in the Spinel Stability Field

Though the amphibolite samples show high LOI values (3.59-6.92 wt. \%), their chondrite-normalized REE and primitive mantle-normalized multi-element spider


Fig. 7. Primitive mantle-normalized multi-element spidergram (A) and chondrite-normalized REE patterns (B). Normalizing values, and OIB, E-MORB, N-MORB trace elemental contents are from Sun and McDonough (1989).
diagrams exhibit smooth and coherent patterns (fig. 7), suggesting the low mobility of the HFSEs, REEs and Y during post-intrusion metamorphism and alteration. Crustal contamination was likely insignificant, as evidenced by: (1) no inherited zircons, (2) no negative $\mathrm{Nb}-\mathrm{Ta}$ anomalies, (3) low REE contents and weakly left-leaning REE patterns, (4) low incompatible element contents and left-leaning primitive man-tle-normalized multi-element patterns, and (5) lack of $\mathrm{Th} / \mathrm{Nb}$ vs. $\mathrm{La} / \mathrm{Sm}, \mathrm{Th} / \mathrm{Nb}$ vs. $\varepsilon \mathrm{Nd}(\mathrm{t})$, or MgO vs. $\mathrm{Nb} /$ La correlations. Mantle-derived primary melts commonly have high Ni ( $>400 \mathrm{ppm}$ ) and $\mathrm{Cr}(>1000 \mathrm{ppm})$ contents, and Mg\# values ( $73-81$ ) (Litvak and Poma, 2010). Our samples show low and variable Mg\#, Cr and Ni contents, indicating olivine and clinopyroxene fractionation. Significant negative Eu anomalies $\left(\mathrm{Eu}^{*}=0.62-0.77\right.$; fig. 7) also indicate plagioclase fractionation.

Our Neoarchean amphibolite samples are geochemically MORB-like (figs. 7A and 7 B ). For the trace elemental ratios that are less affected by partial melting (that is good mantle source tracers), our samples are similar to the average primitive mantle (PM; Sun and McDonough, 1989), for example, $\mathrm{Th} / \mathrm{Nb}=0.11$ ( $\mathrm{PM}=0.12$ ), $\mathrm{Hf} / \mathrm{Sm}=$ $0.65(\mathrm{PM}=0.71), \mathrm{Sm} / \mathrm{Nd}=0.31(\mathrm{PM}=0.32), \mathrm{Lu} / \mathrm{Hf}=0.22(\mathrm{PM}=0.24)$, and $\mathrm{Zr} / \mathrm{Hf}=$ 38 ( $\mathrm{PM}=36$ ). In figure 8, all the samples plot close to the PM , also suggesting strong geochemical resemblance with the PM. No correlations exist between the Mg\# and these trace elemental ratios, suggesting that these ratios reflect the mantle source features, rather than a fractionation artefact. In figure 8 F , the sample $\mathrm{Sm} / \mathrm{Nd}$ ratios (0.30-0.37, avg. 0.32 ) plot below the depleted mantle $\mathrm{Sm} / \mathrm{Nd}$ ratio (0.373) calculated

Table 4
Whole-rock Sm-Nd and Lu-Hf isotopic data of the Biliya amphibolite (Samples D18-1-1 and D18-2-1) and metadiorite (Samples D18-3-1, D18-4-1 D18-5-1 and D18-6-1)

| Sample | $\begin{aligned} & \text { Age } \\ & \text { (Ga) } \\ & \hline \end{aligned}$ | $\frac{{ }^{147} \mathrm{Sm}}{{ }^{144} \mathrm{Nd}}$ | $\frac{{ }^{\frac{143}{} \mathrm{Nd}}}{{ }^{144} \mathrm{Nd}}$ | $\left({ }^{143} \mathrm{Nd} /{ }^{144} \mathrm{~N}\right.$ | $\mathrm{Nd}(\mathrm{t})^{\prime}$ | $\begin{aligned} & \hline \mathrm{T}_{2 \mathrm{DM}} \\ & (\mathrm{Ga}) \\ & \hline \end{aligned}$ | $\frac{{ }^{176} \mathrm{Hf}}{{ }^{177} \mathrm{Hf}}$ | $\frac{{ }^{176} \mathrm{Lu}}{{ }^{177} \mathrm{Hf}}$ | $\left({ }^{176} \mathrm{Hf} /{ }^{177} \mathrm{Hf}\right)_{\mathrm{i}}$ | $\varepsilon_{\text {Hf }}$ <br> (t) | $\begin{aligned} & \mathrm{T}_{2 \mathrm{DM}} \\ & (\mathrm{Ga}) \\ & \hline \end{aligned}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| D18-1-1 | 2.5 | 0.09963 | 0.51122 | 0.50958 | 3.58 | 2.6 | 0.28169 | 0.01030 | 0.28119 | 0.51 | 2.97 |
| D18-2-1 | 2.5 | 0.11757 | 0.51131 | 0.50937 | -0.54 | 2.9 | 0.28187 | 0.01354 | 0.28119 | 1.37 | 2.92 |
| D18-3-1 | 2.5 | 0.19329 | 0.51277 | 0.50958 | 3.67 | 2.8 |  |  |  |  |  |
| D18-4-1 | 2.5 | 0.19826 | 0.51272 | 0.50945 | 1.09 |  |  |  |  |  |  |
| D18-5-1 | 2.5 | 0.19022 | 0.51276 | 0.50963 | 4.51 |  |  |  |  |  |  |
| D18-6-1 | 2.5 | 0.21140 | 0.51279 | 0.50931 | -1.73 | 5.1 | 0.28283 | 0.03562 | 0.28119 | -1.95 | 3.12 |


by Salters and Stracke (2004) at 2500 Ma , also supporting that the amphibolite was extracted from a mantle source with PM geochemical characteristics.

The petrogenesis of the amphibolite is dominated by partial melting processes. Ytterbium is compatible in garnet but La and Sm are not, and thus ( $\mathrm{La} / \mathrm{Yb}$ ) CN and $(\mathrm{Sm} / \mathrm{Yb}) \mathrm{CN}$ would increase during partial melting in the garnet stability field. Sm and Yb have similar partition coefficients in the partial melting of a spinel lherzolite mantle source, and $(\mathrm{Sm} / \mathrm{Yb}) \mathrm{CN}$ ratios are nearly unfractionated in the melting in the spinel stability field, whereas the $\mathrm{Ce} / \mathrm{Sm}$ ratios would decrease (Aldanmaz and others, 2000). Our amphibolite samples exhibit flat REEs patterns with low $\mathrm{La} / \mathrm{Yb}$ ) $\mathrm{CN}(2.0-2.2)$ and $(\mathrm{Sm} / \mathrm{Yb}) \mathrm{CN}(1.4-1.6)$, and they fall into the spinel lherzolite fields (figs. 9A and B). The calculated average pressure and temperature are 2.0 GPa and $1480^{\circ} \mathrm{C}$ (fig. 9C; Till and others, 2012), markedly different from those of present-day Atlantic and Pacific MORB, and correspond to a $\sim 65 \mathrm{~km}$ partial melting depth. To determine the degree of partial melting, we conducted petrogenetic modeling with the continuous melting model devised by Zou and Reid (2001). The results suggest that $\sim 20 \%$ melting of the mantle source with PM-like geochemical features in the spinel stability field could best produce the REEs compositions of our amphibolite samples (fig. 9D and table 5), consistent with the results shown in figures 9A and 9B. Thus, we believed that the lithospheric mantle (that is source rocks of the Biliya amphibolites) show PM-like geochemical features and their $\sim 20 \%$ partial melting produced the Biliya amphibolites.

The question remains as how to interpret the whole-rock Nd and Hf isotopic characteristics. The samples yielded wide-ranged $\varepsilon \mathrm{Nd}(\mathrm{t})(-1.7$ to +4.5$)$ and near-zero $\varepsilon \mathrm{Hf}(\mathrm{t})$ $(-1.9)$ values, which are significantly lower than those of the present-day Indian, Pacific and Atlantic MORBs formed from the depleted mantle (fig. 5B). Recent Rb-Sr, Sm-Nd, Lu-Hf and $\mathrm{Pb}-\mathrm{Pb}$ isotopic (Herzberg and others, 2014; Lambart and others, 2019) and seismic anisotropy studies (Bodmer and others, 2015; Moriwaki and others, 2020) have shown that the Earth's mantle is heterogeneous as a result of early planetary differentiation and subsequent crustal recycling during plate tectonics. Thus, though the Neoarchean lithospheric mantle of the Erguna Terrane have PM-like trace elemental features, its Nd-Hf isotopic compositions are heterogeneous and show a wide range of values.

## Petrogenesis of the Metadiorite: Partial Melting of Mafic Lower Crust

Our Biliya metadiorites show low SiO2 (59.07-64.56 wt.\%) and high Al2O3 (14.53-16.31 wt.\%), and variable Eu anomalies ( $\mathrm{Eu}^{*}=0.77-1.20$ ), suggesting that


Fig. 8. Binary diagrams of the Biliya amphibolite with the PM. Figures A, B, and C are from Dilek and Furnes (2014) and Furnes and others (2015), and figures D, E, and F are from Cook and others (2016). Abbrevations: E-MORB - enriched mid-ocean ridge basalt, BABB - back-arc basin basalt, FAB - forearc basalt, IAT - island arc tholeiite, PM - primitive mantle, DM - depleted mantle.
fractional crystallization did not play an important role in their petrogenesis. The metadiorites show low FeOT/MgO and A/CNK ratios ( $0.84-1.18$ ) ratios, Zr and $\mathrm{Zr}+$ $\mathrm{Nb}+\mathrm{Ce}+\mathrm{Y}$ contents, high whole-rock $\varepsilon \mathrm{Nd}(\mathrm{t})(-0.5-+3.6)$ and $\varepsilon \mathrm{Hf}(\mathrm{t})(+0.5-$ $+1.4)$ values. They contain no high-temperature anhydrous minerals, but some hornblendes. The Biliya metadiorites are I-type granitoids. I-type granite is generally derived from dehydration melting of igneous rock or their metamorphic equivalents (Chappell and others 1998; Chappell 1999). The low A/CNK ratios ( $0.84-1.18$ ) precludes their derivation from melting of a sedimentary source or melting of an igneous source under water-saturated condition because both processes would produce strongly peraluminous melts (Douce and Johnston, 1991; Douce, 1996). The low SiO2 contents (59.07-64.56 wt.\%) for our Biliya metadiorite samples suggest a mafic


Fig. 9. (A) Yb vs. La/ Yb plot (Tschegg and others, 2011), (B) ( $\mathrm{Yb} / \mathrm{Sm}$ ) n vs. (Ce/Sm)n plot (Cook and others, 2016), (C) Pressure-temperature diagram showing the temperature and melting depth of the Biliya amphibolite (Cook and others, 2016; Till and others, 2012), (D) REE modeling results for continuous mantle melting of the PM in the spinel and garnet stability fields. The modeling parameters and data are shown in table 5.
source, such as amphibolite. Experimental studies have demonstrated that partial melting of basaltic rocks can produce intermediate to silicic melts leaving a granulite residue at 8 to 12 kbar or an eclogite residue at 12 to 32 kbar (Rapp and Watson, 1995). Thus, it is likely that our Biliya metadiorites are originated from partial melting of sub-alkaline meta-basalts (that is amphibolite).

Based on the absence or presence respectively of initially inherited zircons, the Itype granites were classified into two distinct types, high- and low-temperature (Chappell and others, 1998). No inherited zircons have been found in our Biliya metadiorite samples. We carried out whole rock zirconium saturation temperature ( TZr ) calculation and got the TZr in a range of 796 to $852^{\circ} \mathrm{C}$, and thus the Biliya metadiorites are high-temperature I-type granitoids. The high-temperature I-type granites formed from a magma that was completely or largely molten (Chappell and others, 1998). Previous experiments reveal that water-unsaturated dehydration melting of the meta-basalts can generate mildly peraluminous melt with high K content (Rapp and Watson 1995; Chappell and others, 2012), and water-saturated melting of the meta-basalts yield strong peraluminous melts (Beard and Lofgren 1991). The Biliya metadiorites are weakly peraluminous ( $\mathrm{A} / \mathrm{CNK}=0.84-1.18$ ), and they should be generated by water-unsaturated dehydration melting of the meta-basalts. The meta-basalts constituted the ancient mafic lower crust of the region.

Experimental data has demonstrated that biotite and muscovite will breakdown when the temperature is higher than 800 to $850^{\circ} \mathrm{C}$ (Thompson and Connolly 1995), but dehydration melting of amphibolites requires much higher temperatures $(>1000$ ${ }^{\circ} \mathrm{C}$, Rapp and Watson, 1995). Residual mineral assembles after melt extraction may also play a crucial role in forming magmas with peculiar geochemical characteristics

Table 5
Trace element modeling results

| Spinel lherzolite partial melting |  |  |  |  |  |  |  | Garnet lherzolite partial melting |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| No. | 1 | , |  | 4 | 5 | 6 | 7 | 1 | , |  |  | 5 |  | 7 |
| F | 0.01 | 0.03 | 0.05 | 0.08 | 0.10 | 0.15 | 0.20 | 0.01 | 0.03 | 0.05 | 0.08 | 0.10 | 0.15 | 0.20 |
| Th | 9.14 | 2.93 | 1.73 | . 07 | 0.86 | 0.57 | 0.43 | 8.68 | 2.93 | 1.73 | 1.07 | 0.86 | 0.57 | 0.4 |
| Nb | 66.19 | 24.50 | 14.54 | 9.02 | 7.19 | 4.78 | 3.58 | 58.86 | 24.28 | 14.53 | 9.02 | 7.19 | 4.78 | 3.58 |
| Ta | 3.81 | 1.41 | 0.84 | 0.52 | 0.41 | 0.27 | 0.21 | 3.38 | 1.40 | 0.84 | 0.52 | 0.41 | 0.27 | 0.21 |
| La | 53.87 | 23.27 | 14.00 | 8.69 | 6.93 | 4.61 | 3.45 | 40.99 | 21.80 | 13.83 | 8.68 | 6.93 | 4.61 | 3.45 |
| Ce | 83.41 | 51.56 | 34.83 | 22.37 | 17.90 | 11.90 | 8.91 | 52.23 | 39.04 | 29.98 | 21.28 | 17.52 | 11.88 | 8.91 |
| Pr | 9.19 | 6.66 | 4.97 | 3.41 | 2.77 | 1.85 | 1.39 | 5.34 | 4.46 | 3.75 | 2.94 | 2.53 | 1.81 | 1.38 |
| Nd | 54.78 | 36.46 | 25.73 | 16.96 | 13.63 | 9.08 | 6.80 | 18.18 | 16.18 | 14.42 | 12.18 | 10.92 | 8.42 | 6.65 |
| Zr | 168.11 | 149.08 | 132.08 | 110.11 | 97.62 | 72.99 | 56.11 | 103.99 | 96.68 | 89.85 | 80.46 | 74.74 | 62.21 | 51.96 |
| H | 4.64 | 4.11 | 3.64 | 3.04 | 2.69 | 2.01 | 1.55 | 2.87 | 2.67 | 2.48 | 2.22 | 2.06 | 1.72 | 1.43 |
| Sm | 6.94 | 6.12 | 5.40 | 4.47 | 3.94 | 2.91 | 2.2 | 3.72 | 3.51 | 3.30 | 3.00 | 2.82 | 2.40 | 2.03 |
| Eu | 2.60 | 2.30 | 2.04 | 1.69 | 1.49 | 1.11 | 0.84 | 1.21 | 1.16 | 1.11 | 1.03 | 0.98 | 0.86 | 0.75 |
| Ti | 11782 | 10996 | 10248 | 9195 | 8541 | 7066 | 5824 | 7220 | 6975 | 6733 | 6373 | 6137 | 5556 | 4994 |
| Gd | 8.59 | 7.70 | 6.88 | . 80 | 5.17 | 3.89 | 2.99 | 3.27 | 3.21 | 3.14 | 3.03 | 2.96 | 2.76 | 2.55 |
| Tb | 1.36 | 1.24 | 1.13 | 0.98 | 0.89 | 0.69 | 0.54 | 0.49 | 0.49 | 0.48 | 0.47 | 0.47 | 0.45 | 0.44 |
| Dy | 9.19 | 8.41 | 7.67 | 6.66 | 6.05 | 4.72 | 3.69 | . 09 | 3.08 | 3.06 | 3.04 | 3.03 | 2.99 | 2.94 |
| Y | 39.05 | 36.98 | 34.95 | 32.02 | 30.14 | 25.71 | 21.69 | 8.89 | 9.15 | 9.45 | 9.98 | 10.41 | 11.90 | 15.39 |
| Ho | 1.88 | 1.74 | 1.60 | 1.41 | 1.29 | 1.03 | 0.82 | 0.60 | 0.60 | 0.61 | 0.61 | 0.61 | 0.62 | 0.63 |
| Er | 4.89 | 4.56 | 4.25 | 3.81 | 3.53 | 2.90 | 2.36 | 1.59 | 1.60 | 1.62 | 1.64 | 1.65 | 1.70 | 1.78 |
| Tm | 0.80 | 0.74 | 0.69 | 0.61 | 0.56 | 0.45 | 0.37 | 0.21 | 0.22 | 0.22 | 0.23 | 0.23 | 0.25 | 0.27 |
| Yb | 4.23 | 4.01 | 3.79 | 3.47 | 3.27 | 2.79 | 2.35 | 1.12 | 1.15 | 1.18 | 1.23 | 1.27 | 1.41 | 1.66 |
| Lu | 0.60 | 0.57 | 0.54 | 0.50 | 0.47 | 0.40 | 0.35 | 0.14 | 0.14 | 0.15 | 0.15 | 0.16 | 0.18 | 0.23 |

(Beard and Lofgren, 1991). The flat HREE patterns and their $\operatorname{TZr}\left(796-852^{\circ} \mathrm{C}\right)$ may suggest amphibole as residual mineral and biotite and muscovite melting. Significant $\mathrm{Sr}\left(\mathrm{Sr}^{*}=0.56-0.73\right)$, and variable Eu anomalies ( $\mathrm{Eu}^{*}=0.77-1.20$ ) imply that plagioclase might also play as residual mineral in the magma chamber (Martin, 1999). Thus, the Biliya metadiorites were generated by dehydration melting of biotite/muscovite from sub-alkaline meta-basalts, leaving amphibole and plagioclase as the major residual minerals.

## Neoarchean Mantle Enrichment and Crustal Extraction in Continental Arc/back-arc Setting

Generation of MORB-type rocks is generally linked to local/regional extension: 1) mantle wedge plume activity (Castro and others, 2010); 2) mid-ocean ridge seafloor spreading (Ridley and others, 2009; Yang and others, 2018); 3) post-orogenic lithospheric delamination (Priestley and others, 2012); and 4) slab rollback-induced back-arc extension (MacLeod and others, 2013). Mantle wedge plume activity is often associated with widespread OIB-type mafic magmatism, and no Neoarchean OIB-type rocks have been found in the Erguna Terrane. Rocks in mid-ocean ridge seafloor spreading show highly depleted Nd and Hf isotopic compositions, with highly positive $\varepsilon \mathrm{Hf}(\mathrm{t})$ and $\varepsilon \mathrm{Nd}(\mathrm{t})$ values (fig. 5B; Kempton and others, 2002). No evidence for major Neoarchean lithospheric delamination has been found in the microcontinents within the CAOB. Thus, the first three scenarios are considered unlikely. In figure 10, our amphibolite samples fall into the back-arc basin basalt (BABB) field or between the volcanic arc basalt (VAB) and MORB fields, and they were thus likely to have formed in a back-arc basin setting. In this setting (fig. 11B), hot asthenosphere upwelling close to the Moho may have provided sufficient heat for the partial melting of lithospheric mantle in the spinel stability field ( $\sim 65 \mathrm{~km}$ ). I-type granites can be generated in various continental tectonic settings, rather than intra-oceanic settings (Roberts


Fig. 10. Tectonic discrimination diagrams of the Biliya amphibolite ( A and B ) and granodiorite ( C and D). Abbrevations: VAT - Volcanic arc tholeiite, E-MORB - enriched mid-oceanic-ridge basalt, NMORB - normal mid-oceanic- ridge basalt, IAT - island-arc tholeiite, VAG - volcanic arc granites, ORG ocean ridge granites, WPG - within-plate granites, SYN-COLG - syn-collisional granites.
and Clemens 1993). In the tectonic discrimination diagrams, the Biliya metadiorites fall into the volcanic-arc granite field (fig. 10). Considering their coexistence, the Biliya amphibolites and metadiorites are likley formed in a continental arc/back-arc setting.

In comparison with present-day Indian, Pacific, and Atlantic MORBs formed from the depleted mantle, our MORB-like amphibolites have enriched Nd-Hf isotopic compositions with a relatively wide range of values $\varepsilon N d(t)=-1.7$ to +4.5 and $\varepsilon H f(t)=$ -1.9 (fig. 5B). These $\varepsilon N d(t)-\varepsilon H f(t)$ values suggest that the Neoarchean lithospheric mantle (that is source of the Biliya amphibolites) has been significantly enriched. Thus, an enriched mantle and TTG crust have been identified in this study. From the lithospheric mantle to the TTG crust, three-stage mantle segregation may have happened: (1) PM-like lithospheric mantle melting: $20 \%$ melt extraction from lithospheric mantle left a depleted mantle and produced the Biliya MORB-type amphibolite (fig. 11B); (2) Mantle enrichment: metasomatism of PM-like lithospheric mantle by subduction-related fluids and/or melts, and substantial arc-type enriched mantle and mafic lower crust were generated (figs. 11A and 11C); and, (3) Lower crust melting: large-scale partial melting of the mafic lower crust produced the TTG crust (fig. 11C), that is crustal extraction.


Fig. 11. Schematic tectonic diagram showing the three-staged mantle segregation and the generations of the PUM, arc-type enriched mantle, and TTG crust.

## CONCLUSIONS

(1) Neoarchean amphibolite and metadiorite are identified in the Erguna Terrane. The former is derived from $\sim 20 \%$ partial melting of the lithospheric mantle in the spinel stability field ( $\sim 65 \mathrm{~km}$ ), and the latter is generated by the partial melting of mafic lower crust.
(2) The Biliya amphibolite and metadiorite were formed in arc/back-arc extension setting, and their identification confirms the presence of Neoarchean basement in the microcontinents within the CAOB.
(3) Three-staged mantle segregation processes have been proposed: (a) $20 \%$ melt extraction from PM-like lithospheric mantle, leaving behind a depleted mantle; (b) subduction-related fluid/melt metasomatism of the PM-like lithospheric mantle
and its partial melting, generating the arc-type enriched mantle and mafic lower crust; (c) partial remelting of the mafic lower crust produced the TTG crust.

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[^0]:    *State Key Laboratory of Petroleum Resources and Prospecting, China University of Petroleum (Beijing), Beijing 102249, China
    ** College of Geosciences, China University of Petroleum (Beijing), Beijing 102249, China
    *** Shenyang Institute of Geology and Mineral Resources/Shenyang Center of Geological Survey, China Geological Survey, Shenyang 110034, China
    ${ }_{\$ 8}^{\$}$ Research Institute of Petroleum Exploration and Development, PetroChina, Beijing 100083, China
    ${ }^{\$ 8}$ Key Laboratory of Orogen and Crust Evolution, Department of Geology, Peking University, Beijing 100871, China
    ${ }^{\$ 8 \$}$ Institute of Mineral Resources, Chinese Academy of Geological Sciences, Beijing 100037, China
    ${ }^{\dagger}$ Corresponding author: Huichuan Liu (liuhuichuan1986@126.com)

